

Propagation Predictions for Newbies (and OTs too)

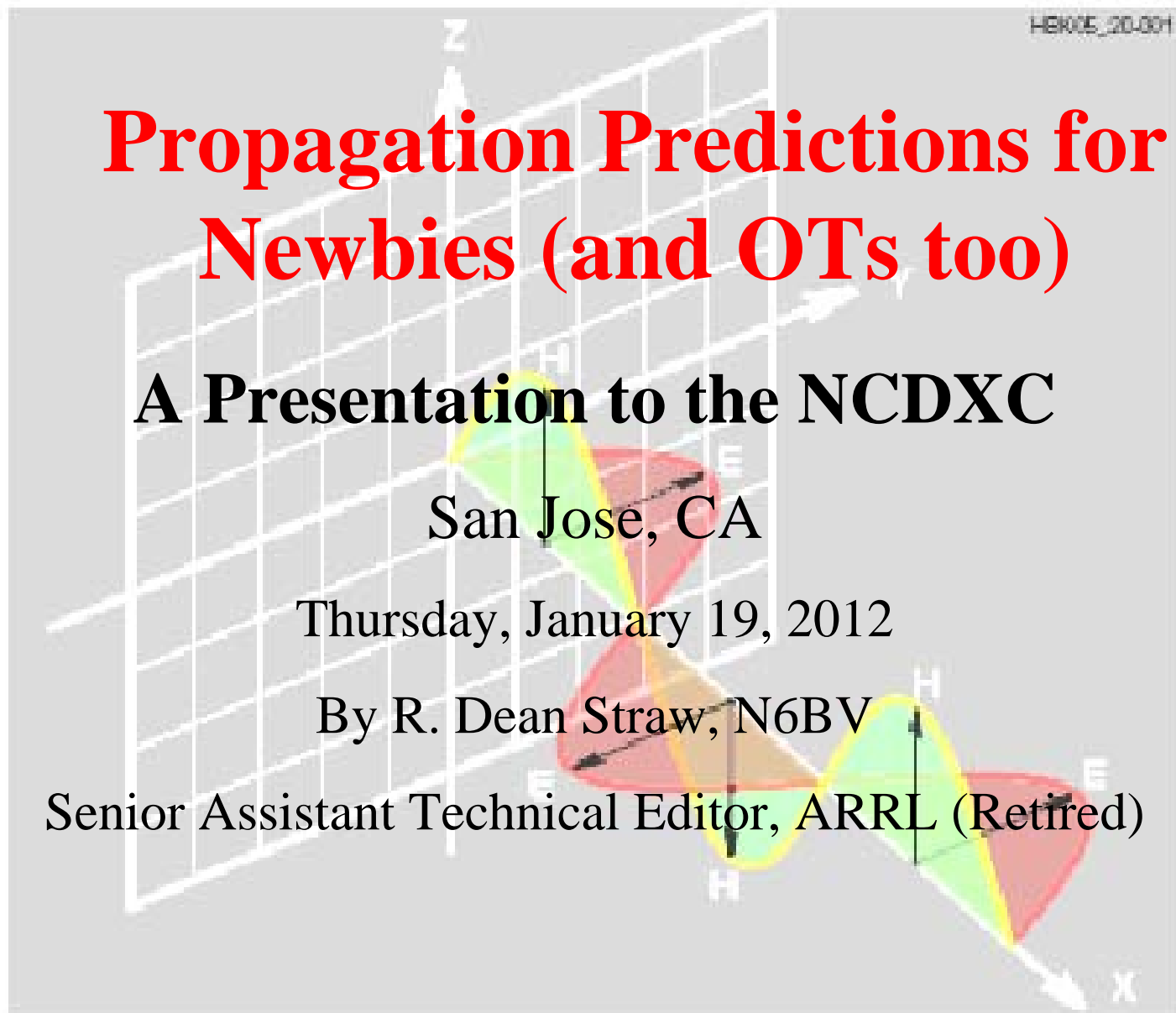
A Presentation to the NCDXC

San Jose, CA

Thursday, January 19, 2012

By R. Dean Straw, N6BV

Senior Assistant Technical Editor, ARRL (Retired)



Albert Einstein — Explaining How Radio Works.

“You see, wire telegraph is a kind of a very, very long cat. You pull his tail in New York and his head is meowing in Los Angeles.”

Albert Einstein — Explaining How Radio Works.

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“Do you understand this?”

Albert Einstein — Explaining How Radio Works.

“You see, wire telegraph is a kind of a very, very long cat. You pull his tail in New York and his head is meowing in Los Angeles.”

“Do you understand this?”

“And radio operates exactly the same way: you send signals here, they receive them there.”

Albert Einstein — Explaining How Radio Works

“The only difference is that there is no cat.”



So How Does a Radio Wave Propagate?

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So How Does a Radio Wave Propagate?

- My usual answer: “It’s magic!”
- However, sometimes I’ll say: “If you wiggle an electron it will produce an electromagnetic field that propagates away from the electron.”
- But all in all, I prefer Einstein’s cat!

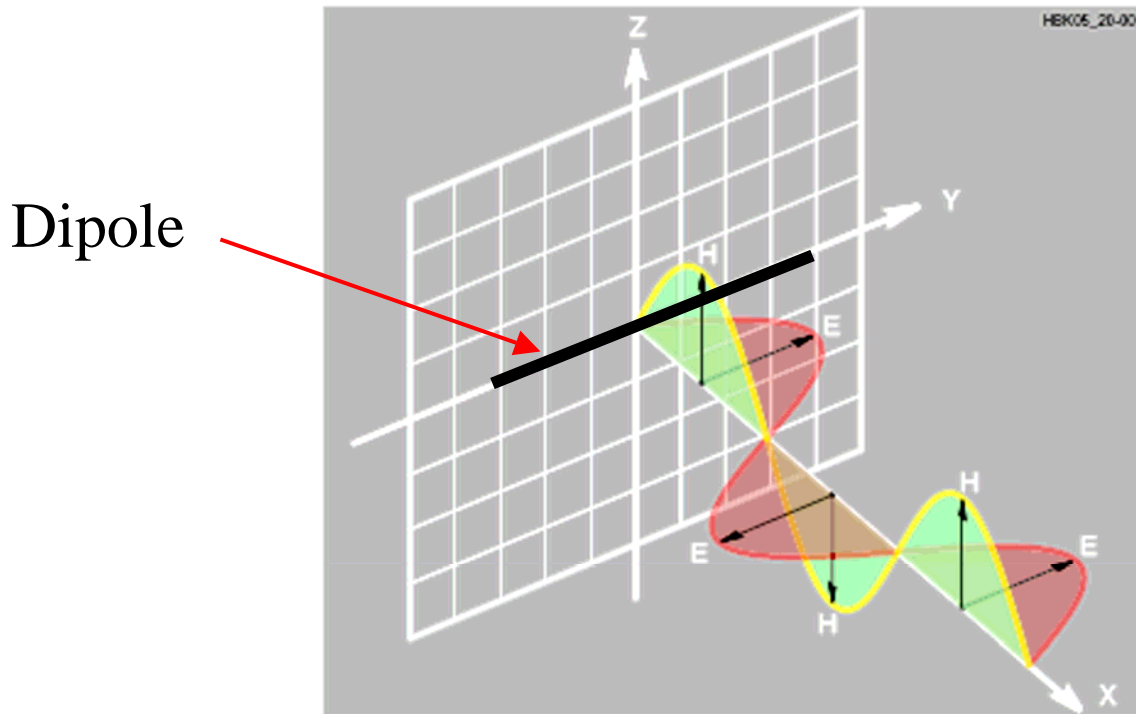


HF Radio Wave Propagation

- Feline humor aside, I'm splitting this talk into two parts:
 - The mechanics of HF propagation (the basics: Propagation 101)
 - Predicting HF propagation (advanced topics: Propagation 201)

The Mechanics of HF Propagation **(the basics: Propagation 101)**

Radio-Wave Propagation



Electric and magnetic field components of an electromagnetic wave. (This is such a cool picture, from *The ARRL Handbook*, 2007 Ed.)

Radio-Wave Propagation

“The Earth is spherical and the waves do not penetrate its surface appreciably, so communication beyond visual distances must be by some means that will bend the waves around the curvature of the Earth.”



(From *The ARRL Antenna Book*, 21st Ed.)

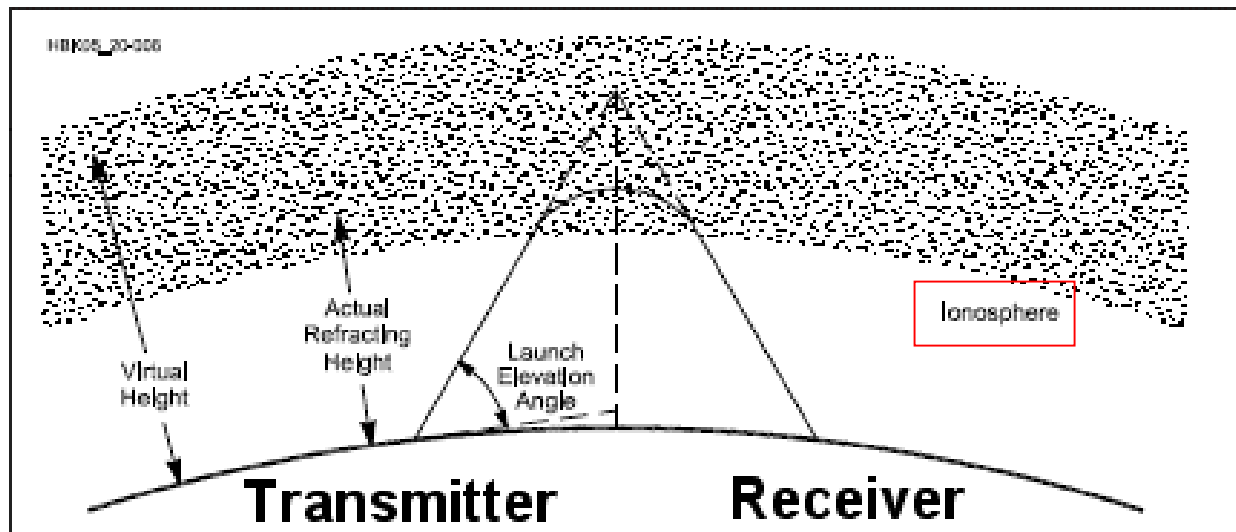
Radio-Wave Propagation

What we need is a “mirror” above the Earth to bounce signals all around the world.

Radio-Wave Propagation

What we need is a “mirror” above the Earth to bounce signals all around the world.

Luckily, we have one. It's called the *ionosphere*.



So, What is the *Ionosphere*?

- The Sun bathes the Earth in a number of different forms of radiation.
 - Visible light.



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 - X-rays.



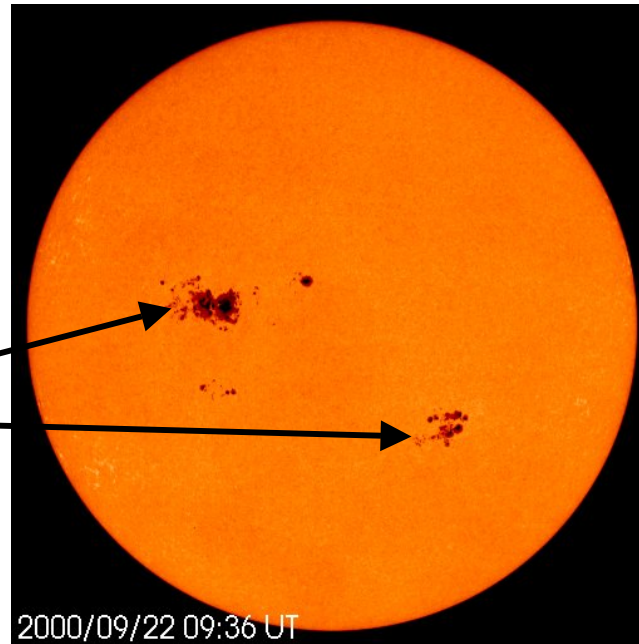
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- The Sun bathes the Earth in a number of different forms of radiation.
 - Visible light.
 - Infrared (heat).
 - X-rays.
 - Ultraviolet (UV) and Extreme UV (EUV).



The Sun's Production of UV

Lots of big
sunspots in
2000



(Courtesy: NASA)

When there are sunspots, the Sun ionizes the Earth's ionosphere strongly with UV, promoting good propagation on the upper HF bands.

What is the *Ionosphere*?

- Radiation from the Sun creates different *regions* in the ionosphere by stripping off electrons from the various types of air molecules there, creating ions.

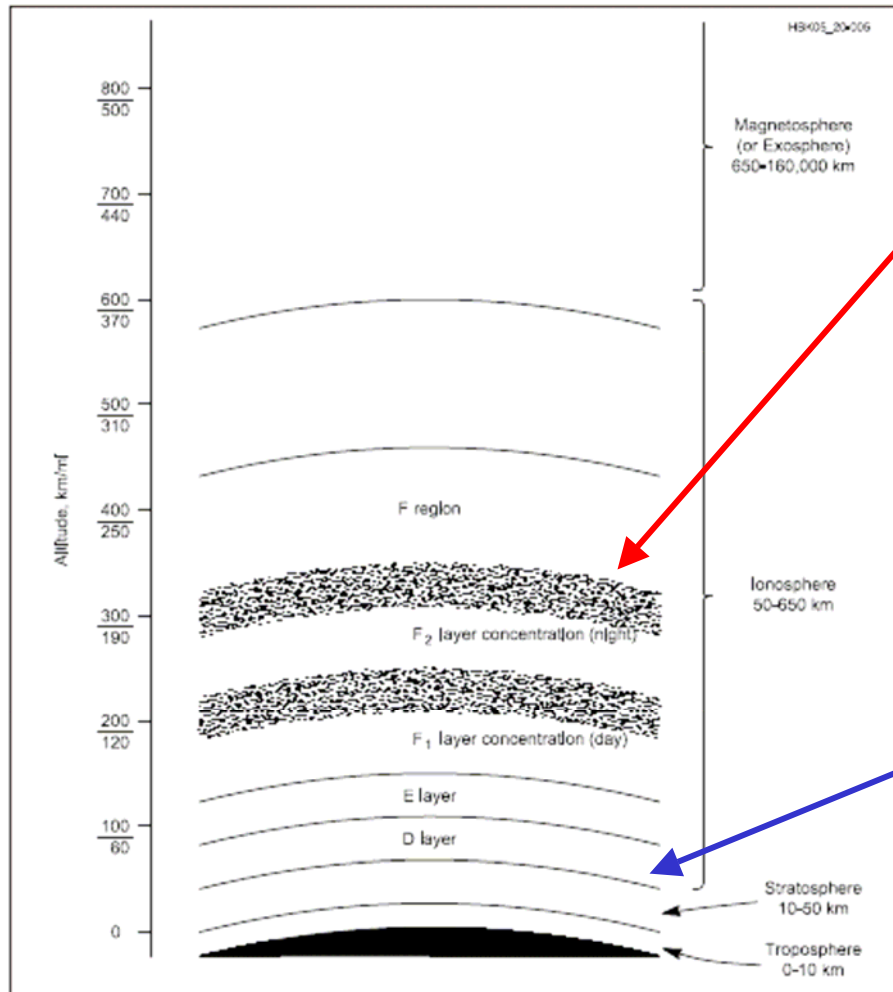
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- Radiation from the Sun creates different *regions* in the ionosphere by stripping off electrons from the various types of air molecules there, creating ions.
- The *free electrons* affect radio waves passing through the ionospheric regions.
- The rate of re-combination of free electrons with positive ions depends on the region.

Regions in the Atmosphere

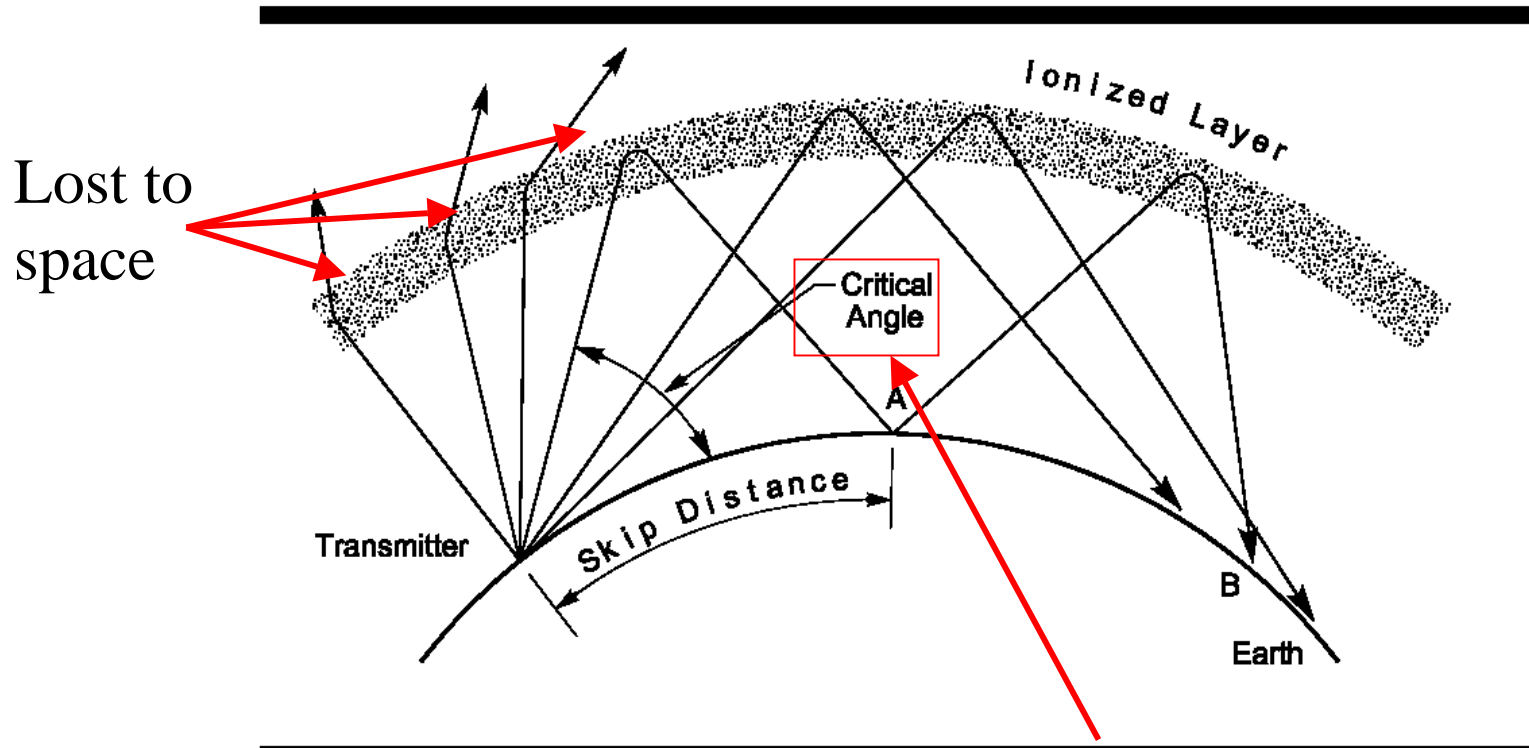


The F₂ region is responsible for most of the long-distance propagation at HF because it's the highest region above Earth and because it has the most free electrons.

The D region is a “bad guy,” absorbing lower-frequency waves

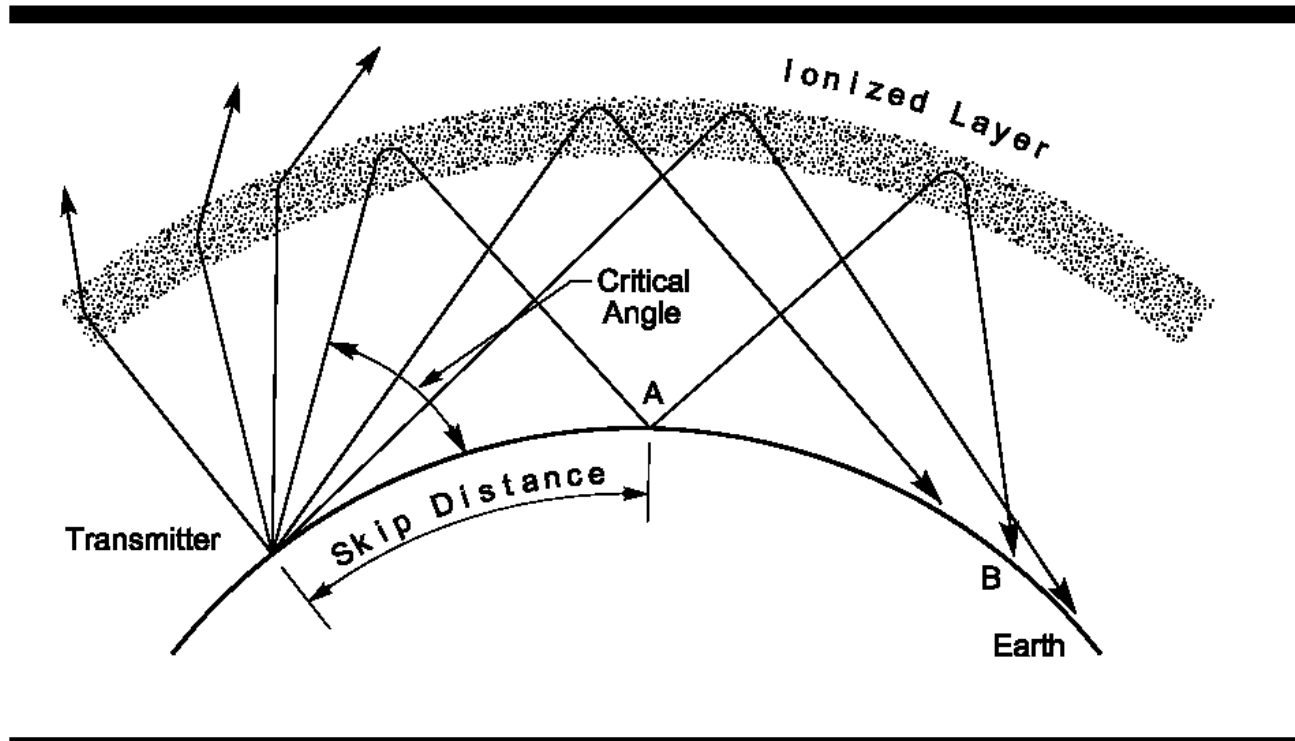
(From *The ARRL Handbook*, 2007 Ed.)

Bending Around the Earth's Curvature



Behavior of waves encountering a simple curved *ionospheric layer* over a curved Earth. Waves higher than the *Critical Angle* are lost to space. Lower-angle waves are refracted and bent back down to Earth. (From *The ARRL Antenna Book*, 21st Ed.)

What About Skip Distance (Zone)?



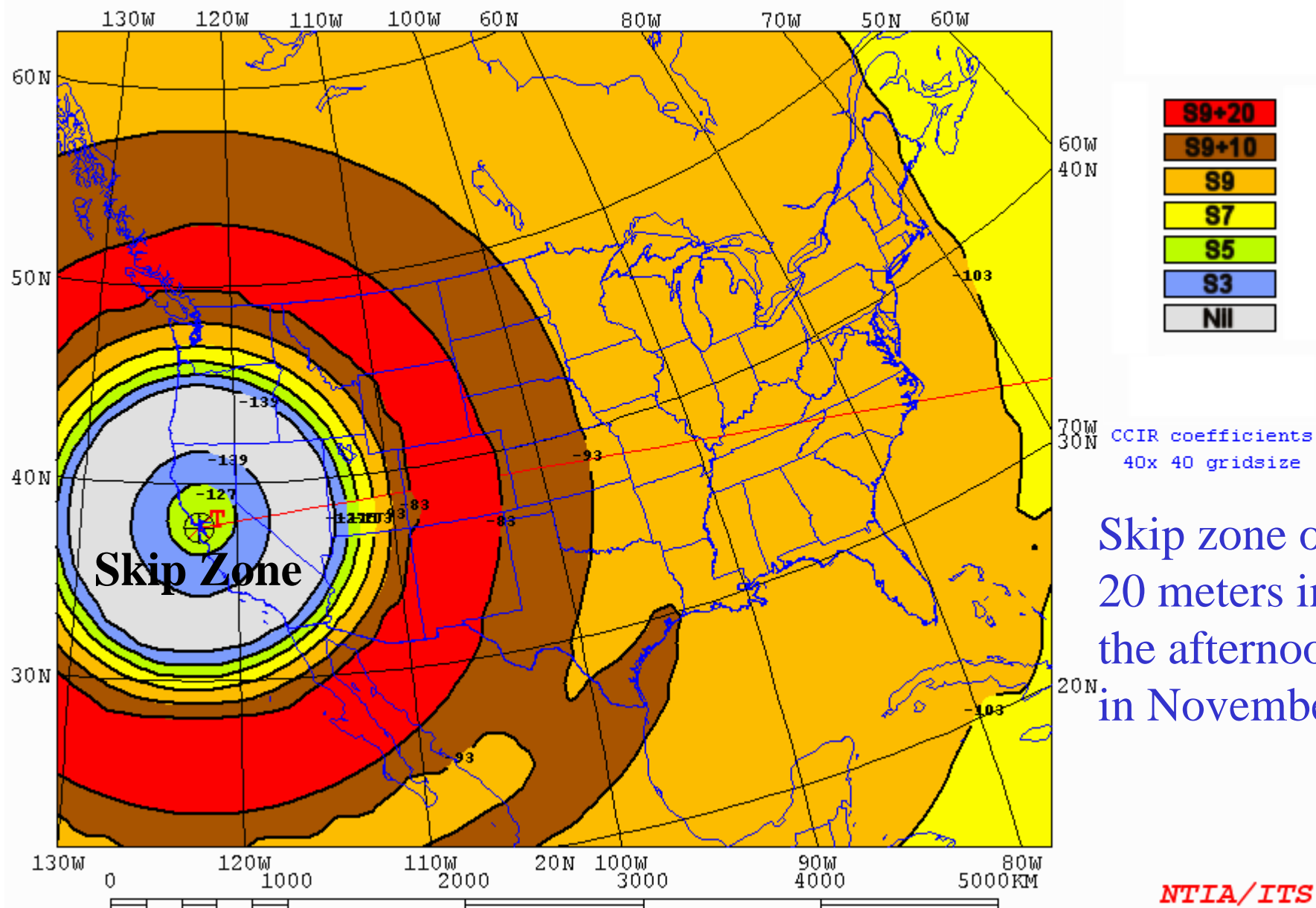
Waves launched at elevation angles between the horizon and the critical angle don't hit the Earth for some distance, after bouncing off the ionosphere. (From *The ARRL Antenna Book*, 21st Ed.)

SAN FRANCISCO [3-el Yagi] 1.5kW 80deg 22ut 14.200MHz Nov 10ssn

SDBW

Tx location to grid of Rx

AREADATA\DEFAULT\SF14.V15



What Influences Propagation Coverage?

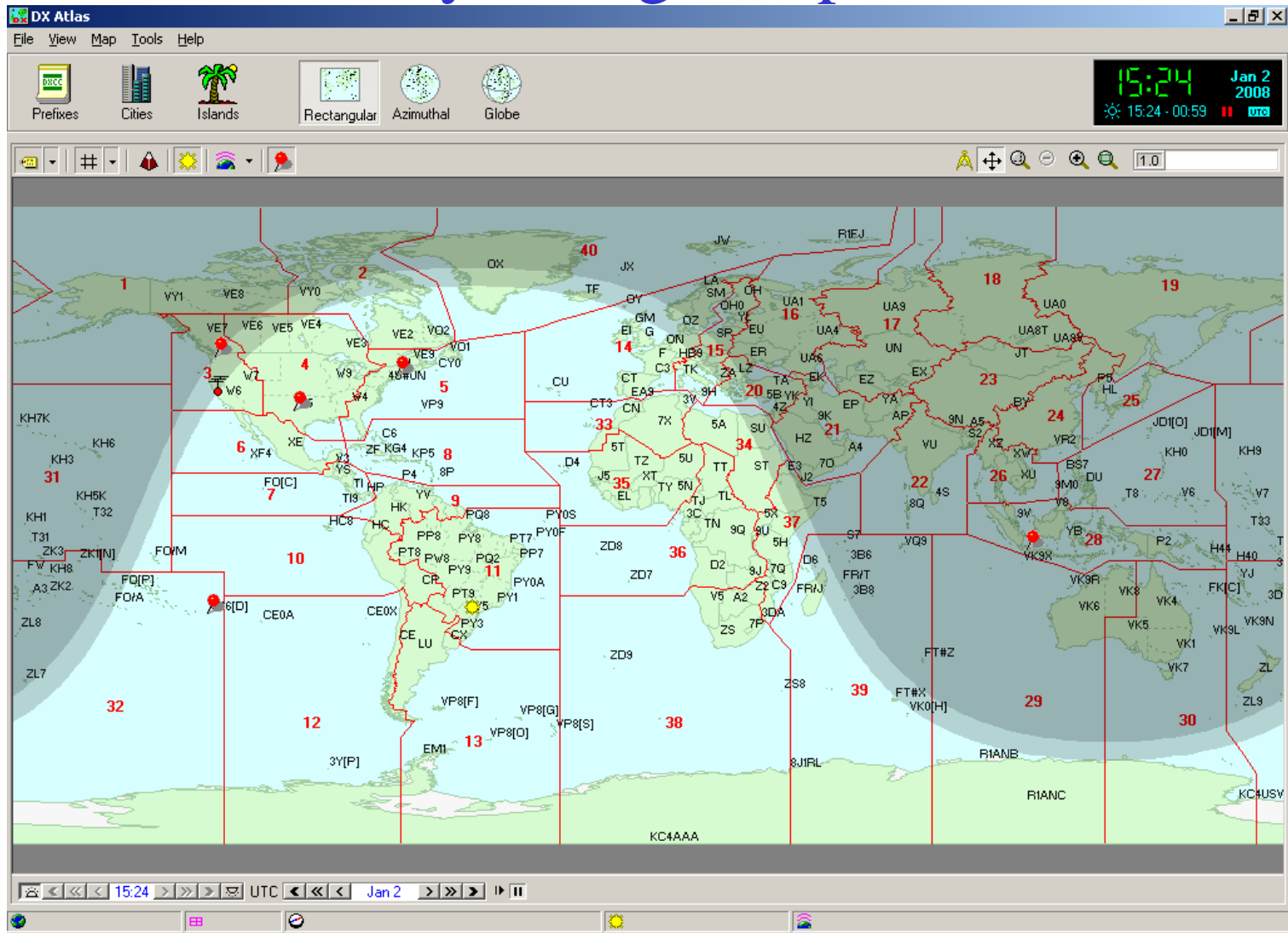
- Frequency, and month/day/hour.
- The state of the ionosphere.
- Antenna gain and transmitter power.
- The launch elevation angle.
- Noise level at receiver (signal-to-noise ratio).

- Frequency, and month/day/hour.
 - 160 to 40 meters (1.8 to 7.3 MHz)
 - Local use during daylight hours
 - Good for DX during the night hours

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 - Transition day-to-night depends on month

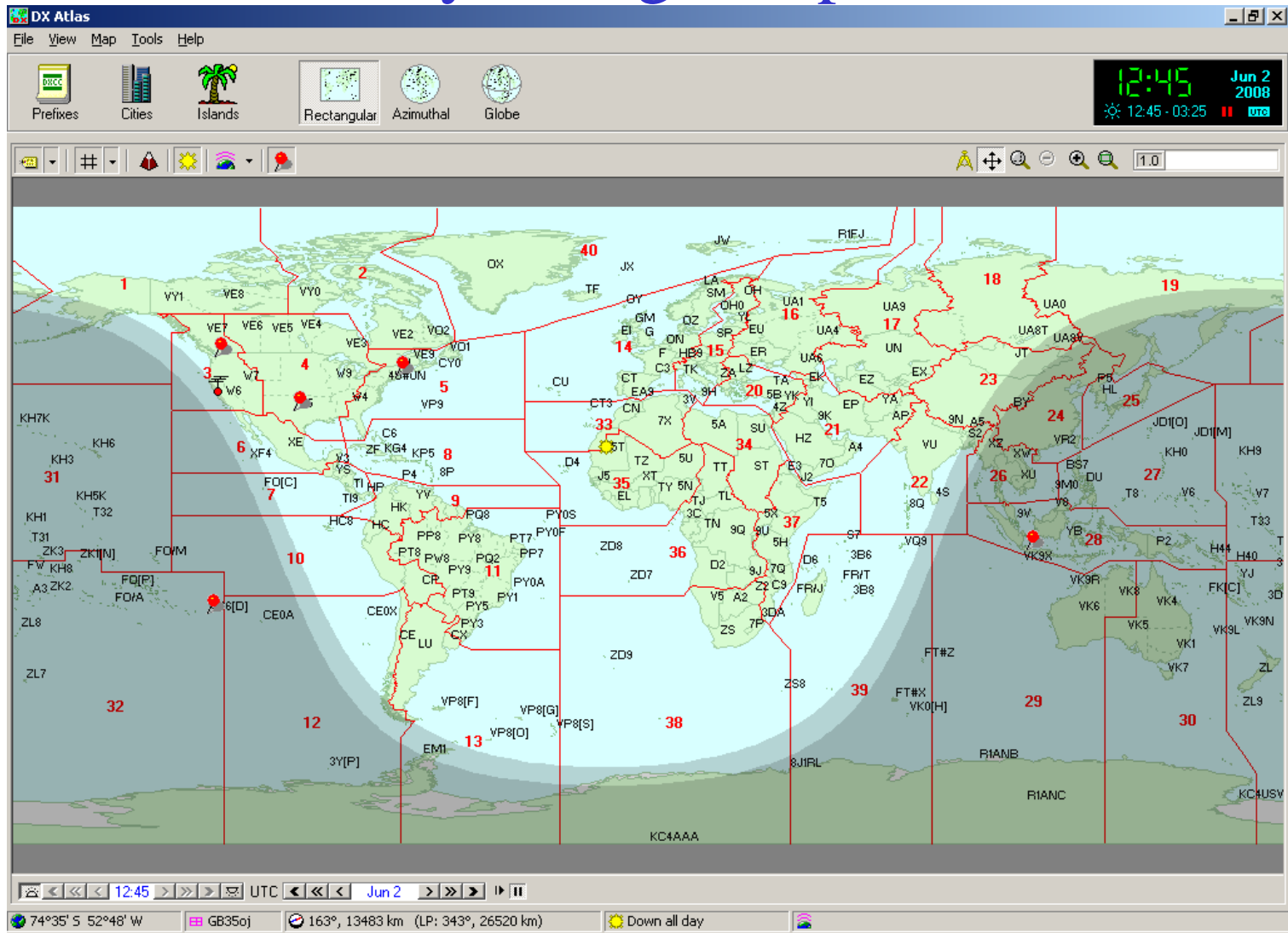
- Transition day-to-night depends on month



Sunrise, W6 in January

DX Atlas

• Transition day-to-night depends on month

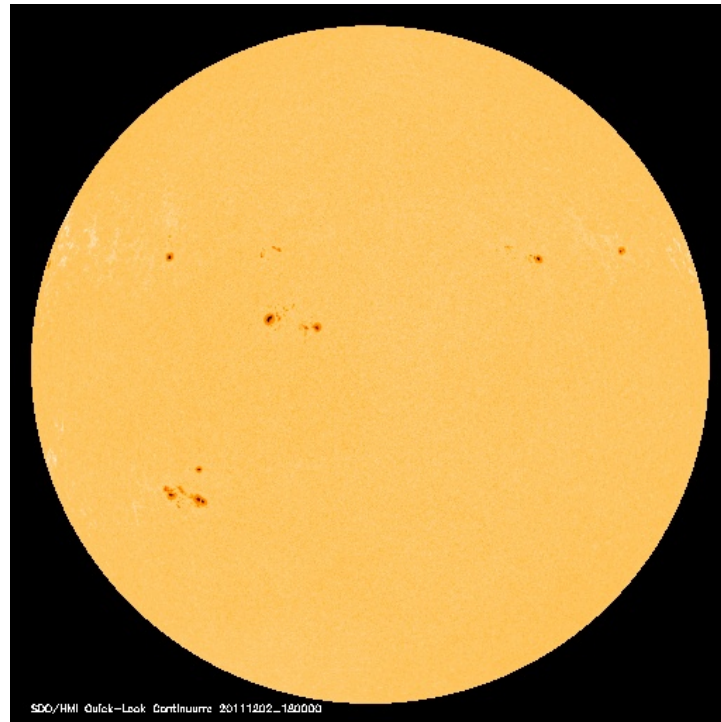


Sunrise, W6 in June

Influences on Propagation Coverage

- Frequency, and month/day/hour.
- The state of the ionosphere.
- Antenna gain and transmitter power.
- The launch elevation angle.
- Noise level at receiver (signal-to-noise ratio).

- The state of the ionosphere.
- The 11-year solar cycle (heading to max.).



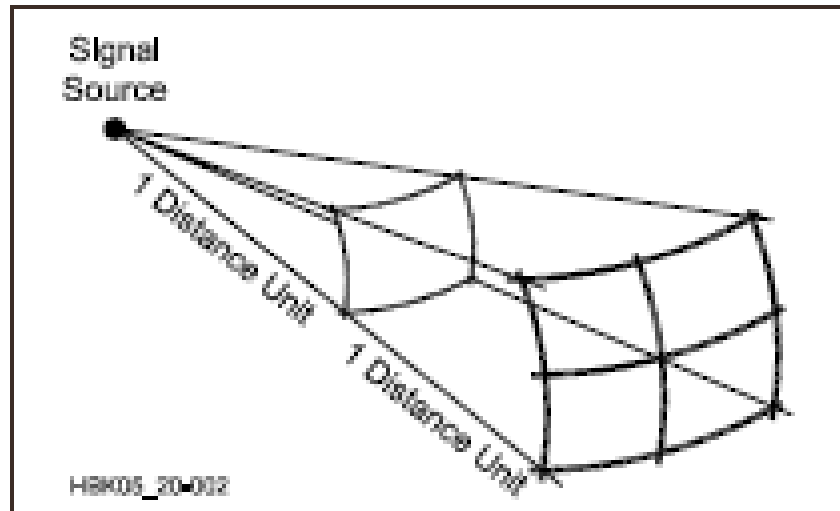
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 - The Earth's geomagnetic field.
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 - The 11-year solar cycle (heading to max.).
 - The Earth's geomagnetic field.
 - Affected by solar flares.
 - Affected by solar Coronal Mass Ejections (CMEs) and the solar wind.

Influences on Propagation Coverage

- Frequency and month/day/hour.
- The state of the ionosphere.
- Antenna gain and transmitter power.
- The launch elevation angle.
- Noise and QRM level at distant receiver (signal-to-noise ratio).

Spreading Out the Energy — the Signal Gets Weaker



Radio energy density “spreads out” as the square of the distance from its source. Therefore, signals get weaker as they depart from your antenna — This is called *spreading loss*. (From *The ARRL Handbook*, 2007 Ed.)

Antenna Gain?

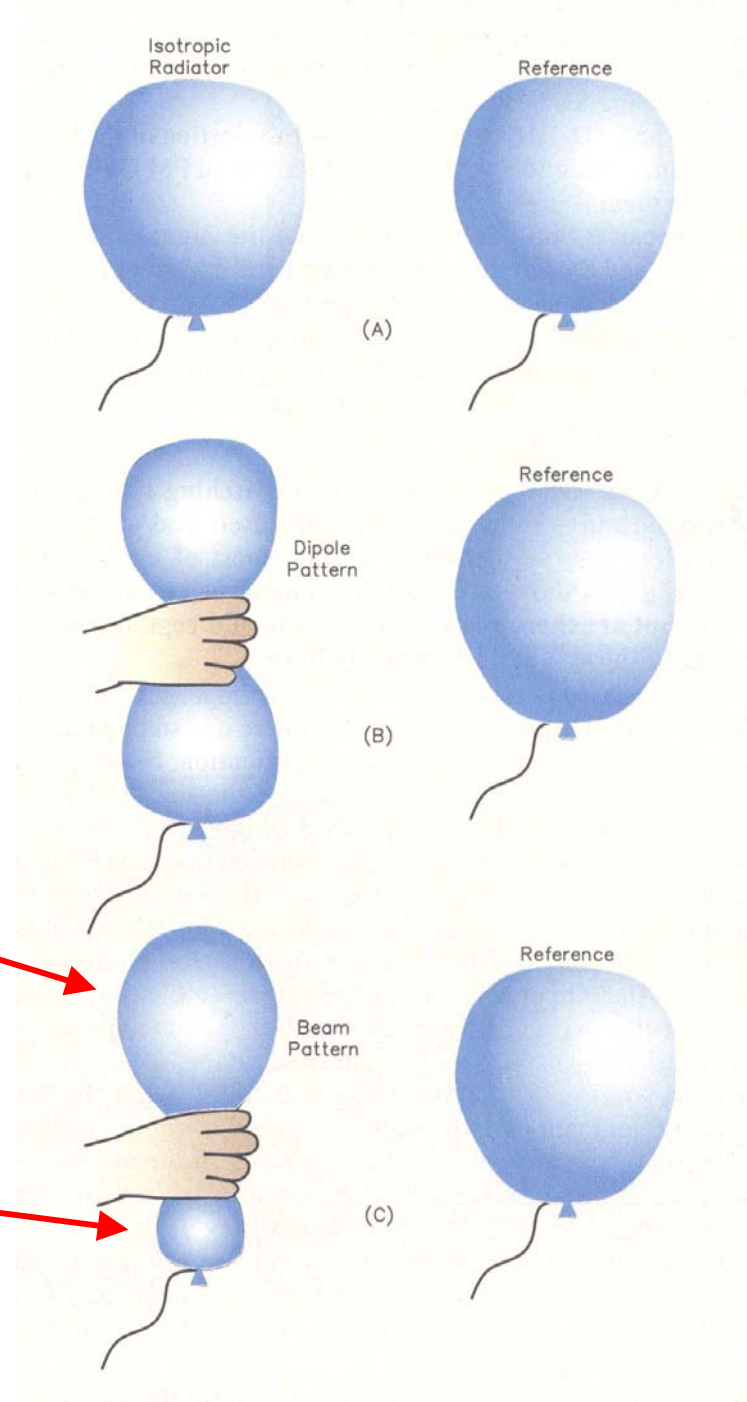
“It is important to remember this so-called *spreading loss* when antenna performance is being considered. Gain can come only from narrowing the radiation pattern of an antenna, which concentrates the radiated energy in the desired direction.”

(From *The ARRL Antenna Book*, 21st Ed.)

Antenna Gain

Like squeezing a balloon...

More signal in one direction means less signal in other directions

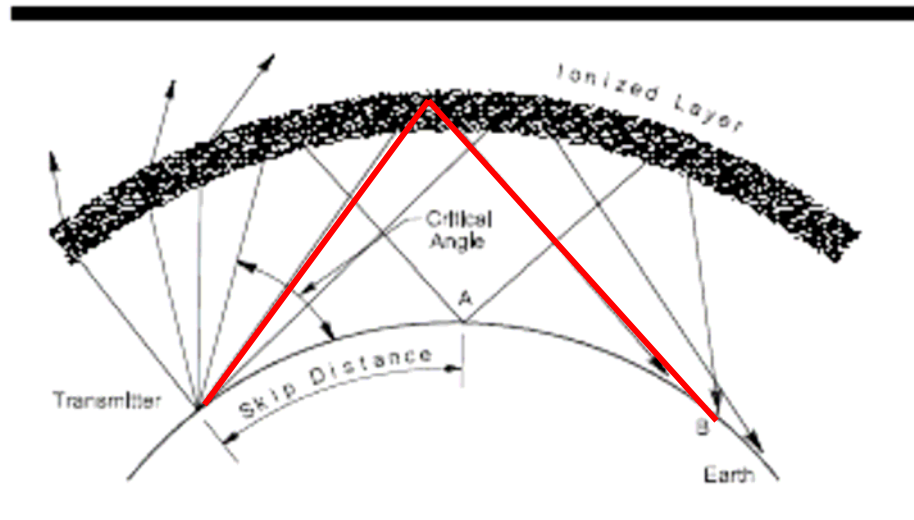


Influences on Propagation Coverage

- Frequency, and month/day/hour.
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- Noise level at receiver (signal-to-noise ratio).

Influences on Propagation Coverage

- The launch elevation angle.
 - Generally speaking, the lower the launch angle from your antenna, the fewer the number of lossy hops necessary to travel to a distant receiver.



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- The launch elevation angle.
 - Generally speaking, the lower the launch angle from your antenna, the fewer the number of lossy hops necessary to travel to a distant receiver.
 - Conversely, to get to a nearby receiver using NVIS (Near Vertical Incidence Skywave) techniques, requires low horizontal antennas that “warm the clouds.”

Influences on Propagation Coverage

- Frequency, and month/day/hour.
- The state of the ionosphere.
- Antenna gain and transmitter power.
- The launch elevation angle.
- Noise level at receiver (signal-to-noise ratio).

Influences on Propagation Coverage

- Noise level at receiver (signal-to-noise ratio).
 - If there is local noise (QRN) or strong interference (QRM) at a distant receiver, your signal can't be heard.

Influences on Propagation Coverage

- Noise level at receiver (signal-to-noise ratio).
 - If there is local noise (QRN) or strong interference (QRM) at a distant receiver, your signal can't be heard.
 - For example, from California you may not be able to work a station in Thailand because stations in Japan or Europe may be much louder than you are.

Predicting HF Propagation

(advanced topics: Propagation 201)

Predicting HF Propagation (advanced topics: Propagation 201)

- The following prediction tables come from the *N6BV Propagation* tables.

N6BV Propagation Predictions

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Detailed Propagation Tables

USA W1B Boston, MA W2A Albany, NY W2N Buffalo, NY W3D Washington, DC W3P Pittsburg, PA W4A Montgomery, AL W4F Miami, FL W4G Atlanta, GA W4K Louisville, KY W4N Raleigh, NC W4S Columbia, SC W4T Memphis, TN W5A Little Rock, AR W5H Houston, TX W5L New Orleans, LA W5M Jackson, MS W5N Albuquerque, NM W5O Oklahoma City, OK W5T Dallas, TX W6L Los Angeles, CA W6S San Francisco, CA W7A Phoenix, AZ W7I Boise, ID W7M Helena, MT W7N Las Vegas, NV W7O Portland, OR W7U Salt Lake City, UT W7W Seattle, WA W7Y Cheyenne, WY W8M Detroit, MI W8O Cincinnati, OH W8W Charleston, WV W9C Chicago, IL W9I Indianapolis, IN W9W Milwaukee, WI W0C Denver, CO W0D Bismarck, ND W0I Kansas City, MO W0K Middle of US, KS W0M St. Louis, MO W0N Omaha, NE	Other, North America 6Y Kingston, Jamaica 8P Bridgetown, Barbados C6A Nassau, Bahamas FG Guadeloupe HP Panama City, Panama HR Tegucigalpa, Honduras J3 Grenada J7 Dominica KL7 Anchorage, Alaska KP2 Virgin Islands OX Godthaab, Greenland TI San Jose, Costa Rica V3 Belmopan, Belize VE1 Halifax, Nova Scotia VE2 Montreal, Quebec VE3 Toronto, Ontario VE4 Winnipeg, Manitoba VE5 Regina, Saskatchewan VE6 Calgary, Alberta VE7 Vancouver, BC VE8 Yellowknife, NWT VO1 St. John's, NFL VP2 Anguilla VP5 Turks & Caicos VP9 Bermuda XE1 Mexico City, Mexico YV0 Aves Island ZF Cayman Islands	South America 8R Georgetown, Guyana CE Santiago, Chile CP La Paz, Bolivia FY Cayenne, French Guiana HC Quito, Ecuador HC8 Galapagos Islands HK Bogota, Columbia LU Buenos Aires, Argentina OA Lima, Peru P4 Aruba PY1 Rio de Janeiro, Brazil PY5 Porto Alegre, Brazil PY0 Fernando de Noronha VP8 Falkland Islands YV Caracas, Venezuela ZP Asuncion, Paraguay	Asia 1S Spratly Islands 3W Ho Chi Minh City, Vietnam 4J Baku, Azerbaijan 4L Tbilisi, Georgia 4S Colombo, Sri Lanka 4X Jerusalem, Israel 5B Nicosia, Cyprus 7O Aden, So. Yemen 9K Kuwait City, Kuwait 9N Katmandu, Nepal A5 Thimbu, Bhutan A7 Ad-Dawhah, Qatar AP Karachi, Pakistan BV Taipei, Taiwan BY1 Beijing, China BY4 Shanghai, China BY0 Lhasa, China EP Tehran, Iran EX Frunze, Kirghiz EY Samarkand, Tadzhik EZ Ashkhabad, Turkoman HL Seoul, Korea HS Bangkok, Thailand HZ Riyadh, Saudi Arabia JA1 Tokyo, Japan JA3 Osaka, Japan JA8 Sapporo, Japan JT Ulan Bator, Mongolia TA Ankara, Turkey UA9 Perm, Russia UA0 Khabarovsk, Russia UN Alma-Ata, Kazakh VR2 Hong Kong VU New Delhi, India VU4 Andaman Islands VU7 Laccadive & Nicobar Islands XZ Rangoon, Myanmar YA Kabul, Afghanistan YI Baghdad, Iraq YK Damascus, Syria
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4 of 11

240+ possible QTHs around the world

N6BV Propagation Predictions

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Detailed Propagation Tables for CA (San Francisco)

SSN Ultra High	SSN Very High	SSN High	SSN Medium	SSN Low	SSN Very Low
January	January	January	January	January	January
February	February	February	February	February	February
March	March	March	March	March	March
April	April	April	April	April	April
May	May	May	May	May	May
June	June	June	June	June	June
July	July	July	July	July	July
August	August	August	August	August	August
September	September	September	September	September	September
October	October	October	October	October	October
November	November	November	November	November	November
December	December	December	December	December	December

Each of the month names in the above table is a hyperlink to the corresponding page of this document. Click on the one you want to jump to.

1 of 649

Choose month, level of solar activity
(<http://www.nwra.com/spawx/ssne.html>)

N6BV Propagation Predictions

20 Meters: Jan., CA (San Francisco), for SSN = High, Sigs in S-Units. (c) 2010 Dean Straw, N6BV

Zone	UTC -->	00	01	02	03	04	05	06	07	08	09	10	11	12	13	14	15	16	17	18	19	20	21	22	23
KL7 = 01		9+	9+	9+	9+	9+	7	2	-	-	1	5	2	-	-	-	1	7	9+	9+	9+	9+	9+	9+	9+
VO2 = 02		9+	9+	7	1	-	-	-	-	-	-	-	-	-	-	-	9	9	9	9	9	9	9	9+	9+
W6 = 03		5	6	6	7	8	8	8	8	8	8	8	8	8	8	8	7	7	5	5	7	8	8	8	6
W9 = 04		9+	9+	9+	9+	9	5	3	5	7	8	4	-	-	-	4	9+	9+	9+	9+	9	9+	9+	9+	9+
W3 = 05		9+	9+	8	2	9	8	7	9	9	9	8	1	-	6	2	9+	9	9	9	9	9	9	9+	9+
XE1 = 06		9+	9+	9+	9+	9+	9+	9	9	9+	9+	9+	9	6	8	9+	9+	9+	9	9	8	8	9	9	9+
TI = 07		9+	9+	9+	9+	8	5	2	4	6	7	2	-	-	1	9+	9	8	7	2	3	5	7	9	9+
VP2 = 08		9+	9+	9	8	6	2	1	3	6	6	2	-	-	1	9	8	7	4	3	5	7	8	9	9+
P4 = 09		9+	9+	9+	9	8	4	2	4	7	7	2	-	-	2	9	8	7	4	2	4	4	8	9	9+
HC = 10		9+	9+	9+	9+	9	8	6	7	9	7	1	-	-	4	9	8	5	2	1	1	2	5	8	9
PY1 = 11		9	9	9	9	9	9	8	8	9	9	8	5	1	2	1	-	-	-	-	-	-	2	5	8
CE = 12		9	9	9	9	9	9	9	9	9	9	8	6	1	3	7	2	1	-	-	-	-	2	4	8
LU = 13		8	9	9	9	9	9	9	9	9	9	8	5	1	4	3	1	-	-	-	-	-	3	4	7
G = 14		2	3	3	4	2	-	1*	-	-	1	-	-	-	-	-	8	8	8	8	7	4	1	-	1
I = 15		2	2	3	3	1	1*	1*	1	5	4	-	-	-	-	-	8	8	6	8	5	1	-	1	1
UA3 = 16		7	6	6	6	5	-	-	2*	1	3	-	-	-	-	-	1	8	6	4	3	2	4	5	5
UN = 17		2	5	8	8	6	-	-	-	-	-	2	-	-	-	1*	-	3	5	5	6	5	4	2	1
UA9 = 18		6	8	9	8	6	-	-	-	-	1*	1	-	-	-	1*	-	-	6	5	4	2	1	-	1
UA0 = 19		9	9	9	9	7	6	4	-	-	-	1	1	1	-	-	-	-	1	-	-	8	8	8	8
4X = 20		6	6	5	4	2	-	-	1	1	1	-	-	-	-	-	8	8	8	6	4	4	4	5	6
HZ = 21		6	4	2	6	3	1*	-	-	-	-	-	-	-	-	1*	4*	8	5	4	2	5	4	5	6
VU = 22		5	7	7	7	3	-	-	-	-	-	-	1	-	-	1*	1*	5	2	2	4	4	4	3	3
JT = 23		8	9	9	8	7	2	-	-	-	-	1	-	4	-	1*	-	-	6	5	2	1	2	4	8
VR2 = 24		8	6	5	5	5	4	2	-	-	1*	1*	1	1	-	-	-	-	4	4	-	-	-	2	6
JA1 = 25		7	7	8	8	8	6	5	-	-	-	-	4	2	-	-	-	-	-	-	-	4	8	7	6
HS = 26		5	6	5	3	1	-	-	-	-	-	1*	1*	-	1*	1*	1*	1*	8	6	2	-	-	-	5
DU = 27		4	3	2	4	4	5	4	1	-	1*	1*	-	1	-	-	-	-	9	8	-	-	1	5	4
YB = 28		8	3*	1*	-	-	-	-	-	-	-	2*	-	1	1*	-	-	1*	9	8	7	5	1	3	2
VK6 = 29		2*	1*	-	-	-	-	1	2	2	1	1	5	2	1*	1*	7	8	7	5	3	1	2*	2*	2*
VK3 = 30		-	-	-	1	4	5	7	7	7	5	4	6	8	5	2*	1*	9	8	6	4	-	-	1*	-
KH6 = 31		9	9+	9+	9+	9+	8	4	9+	9	8	7	9	9	8	2	4	9+	9+	9+	9	9	8	8	9
KH8 = 32		5	6	8	9	9	9	9	9	9	8	6	8	8	7	2	1	7	9	8	6	5	4	3	3
CN = 33		3	1	1	4	5	1	1*	1*	5	5	-	-	-	-	3	7	9	9	7	7	8	7	7	5
SU = 34		7	7	6	5	5	-	1*	1	2	2	-	-	-	-	-	8	8	5	7	5	4	5	5	6
6W = 35		9	9	9	9	8	5	4	1*	5	8	1	-	-	-	2	2	2	4	4	5	6	8	8	9
D2 = 36		9	9	9	8	6	3	2	3	4	1	-	-	-	-	-	2	8	4	5	5	5	6	7	8
5Z = 37		8	8	8	4*	5	1*	-	-	-	-	-	-	-	-	-	6	6	6	4	2	5	6	7	8
ZS6 = 38		9	9	9	9	7	2	1	1	-	-	-	-	-	-	-	7	6*	4*	2	2	4	6	8	8
FR = 39		8	8	8	8	2	1*	-	-	-	-	-	1*	1*	-	2*	6*	8	8	8	4	4	5	5	8
EJL = 40		6	7	8	9	1	-	-	-	-	-	-	-	-	-	-	-	6	6	8	8	8	7	6	6

USA

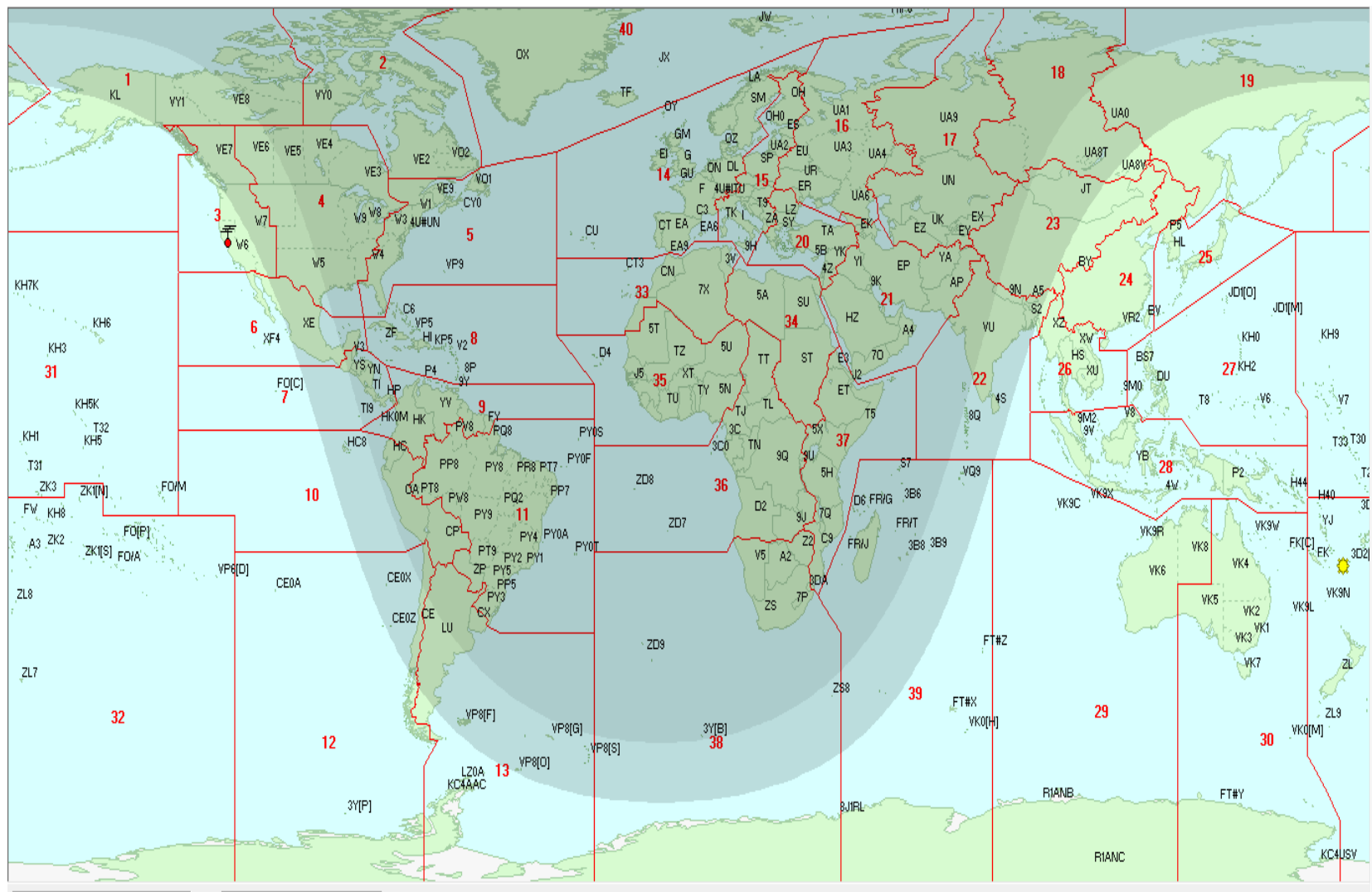
EU

JA

Zone UTC --> * = Longpath
Expected signal levels using 1500 W and 12 dBi isotropic antennas.

Detailed pages for 160/80/40/30/20/17/15/12/10 meters

CQ Zones



New Detailed Prediction Tables

30 Meters: Oct., CA (San Francisco), for SSN = Low, Sigs in S-Units. By N6BV, ARRL.

Zone	UTC -->	00	01	02	03	04	05	06	07	08	09	10	11	12	13	14	15	16	17	18	19	20	21	22	23
KL7 = 01		9	9	9	9+	9+	9+	9	9	9	9	9	9	9	8	9	9+	9	9	9	8	8	8	8	9
VO2 = 02		8	8	8	7	5	4	4	4	3	-	-	-	3	8	8	6	5	3	2	3	5	6	7	8
W6 = 03		9	8	8	5	5	5	5	5	5	5	5	5	5	5	5	5	8	9	8	9	9	9	9	9
W9 = 04		9+	9+	9+	9+	9+	9	9+	9+	9+	9+	9	9	9	9+	9+	9	8	7	5	5	7	8	8	9
W3 = 05		9	9	9	6	4	5	5	5	6	4	9	9	2	9	8	8	7	5	4	4	5	7	8	8
XB1 = 06		9	9	9+	9+	9+	9+	9+	9+	9+	9+	9+	9+	9+	9+	9	8	6	5	2	4	5	7	8	
TI = 07		8	9	9	9	9	9	9	9	9	9	9	6	9	9	8	5	2	1	-	-	1	5	7	
VP2 = 08		9	9	9	8	9	9	9	9	9	8	8	8	8	8	5	1	1	-	-	-	1	6	7	
P4 = 09		8	9	9	9	9	9	9	9	9	8	8	8	8	8	6	2	1	-	-	-	2	6	7	
HC = 10		8	8	9	9	9	9	9	9	9	8	5	4	8	8	5	2	-	-	-	-	-	2	7	
PY1 = 11		5	7	7	7	8	8	7	8	7	6	5	5	1	-	-	-	-	-	-	-	-	1	2	
CE = 12		5	7	8	8	8	8	8	9	9	8	8	6	5	1	-	-	-	-	-	-	-	-	2	
LU = 13		5	7	7	8	8	8	8	8	8	7	7	7	5	1	-	-	-	-	-	-	-	-	2	
G = 14		6	6	7	7	5	5	5	5	6	5	3	1	2	4	3	5	2	1	1	1	2	5	6	
I = 15		6	7	7	6	5	5	7	7	6	5	2	1	2	2	2	4	1	1	-	1	2	4	5	
UA3 = 16		5	5	5	5	6	6	6	5	4	4	3	2	2	5	6	5	4	4	2	2	3	4	5	
UN = 17		4	2	3	3	1	2	2	1	1	2	3	4	4	4	6	7	7	5	4	2	1	1	2	
UA9 = 18		4	7	5	5	4	4	4	2	2	3	4	3	3	4	6	6	5	4	2	1	1	1	2	
UA0 = 19		3	3	4	5	6	7	7	7	7	8	8	8	8	7	7	8	8	7	6	4	3	2	1	
4X = 20		6	6	6	6	5	5	4	3	2	1	1	1	1	2	2	5	2	1	1	1	-	1	3	
HZ = 21		5	5	4	4	5	2	1	-	-	-	-	1	3	4	4	2	2	1	-	1	1	2	4	
VU = 22		2	2	2	1	1	-	-	-	-	1	4	4	5	5	5	6	3	2	1	-	1	1	2	
JT = 23		2	3	4	4	3	3	4	4	4	5	5	5	5	6	8	7	6	5	3	2	2	2	2	
VR2 = 24		-	-	-	-	1	2	4	5	6	8	8	8	8	7	8	8	7	5	4	2	1	-	1	
JA1 = 25		1	1	2	3	5	6	7	6	7	7	7	8	8	6	5	8	8	7	6	5	2	1	-	
HS = 26		-	-	-	-	-	-	-	2	3	5	6	6	6	6	7	6	5	4	2	1	-	-	-	
DU = 27		-	-	-	-	-	2	5	6	7	8	8	8	8	8	8	8	7	5	3	1	-	-	-	
YB = 28	1*	-	-	-	-	-	-	-	3	5	7	8	8	8	8	7	7	6	5	3	1	-	-	-	
VK6 = 29	-	-	-	-	-	-	-	1	4	6	7	8	8	8	8	8	7	6	2	1	-	-	-	-	
VK3 = 30	-	-	-	-	-	2	4	7	8	8	8	9	9	8	8	8	7	6	3	1	-	-	-	-	
KH6 = 31	7	8	9	9	9+	9+	9	8	7	7	8	9	9	9	7	5	9	9	9	8	6	3	2	2	
KH8 = 32	-	1	3	6	8	8	9	9	9	9	9	9	9	9	8	7	7	5	2	-	-	-	-	-	
CN = 33	6	7	8	8	8	7	5	7	7	6	3	1	1	1	1	1	-	-	1	1	1	1	2	6	
SU = 34	7	6	6	6	6	5	5	3	2	1	1	-	1	2	1	6	1	2	1	-	-	1	3	5	
6W = 35	7	8	8	8	8	8	8	8	8	7	5	2	1	-	-	-	-	-	-	-	-	1	4	6	
D2 = 36	5	6	6	7	7	6	5	4	1	-	-	-	-	-	-	-	-	-	-	-	-	-	1	3	
5Z = 37	4	5	4	5	5	3	1	-	-	-	-	-	-	-	1*	2*	6	1*	1	1	-	-	1	2	
ZS6 = 38	5	6	6	7	7	5	3	1	-	-	-	-	-	-	1*	1*	2*	-	-	-	-	-	-	2	
FR = 39	5	3	2	1	1	-	-	-	-	-	-	-	-	-	2	2	2	3	1	-	-	-	1	1	
FJL = 40	6	6	5	6	7	5	3	2	2	2	2	1	1	2	7	7	6	6	5	5	2	5	4	5	
Zone	UTC -->	00	01	02	03	04	05	06	07	08	09	10	11	12	13	14	15	16	17	18	19	20	21	22	23

* = Longpath

Expected signal levels using 1500 W and 6 dBi isotropic antennas.

Example of new 30-meter table

Visualization — Area Coverage Maps

The following hand-made area-coverage maps are for San Francisco, CA, during a period of low solar activity in November (think ARRL Sweepstakes contest).

Visualization — Area Coverage Maps

The following hand-made area-coverage maps are for San Francisco, CA, during a period of low solar activity in November (think ARRL Sweepstakes contest).

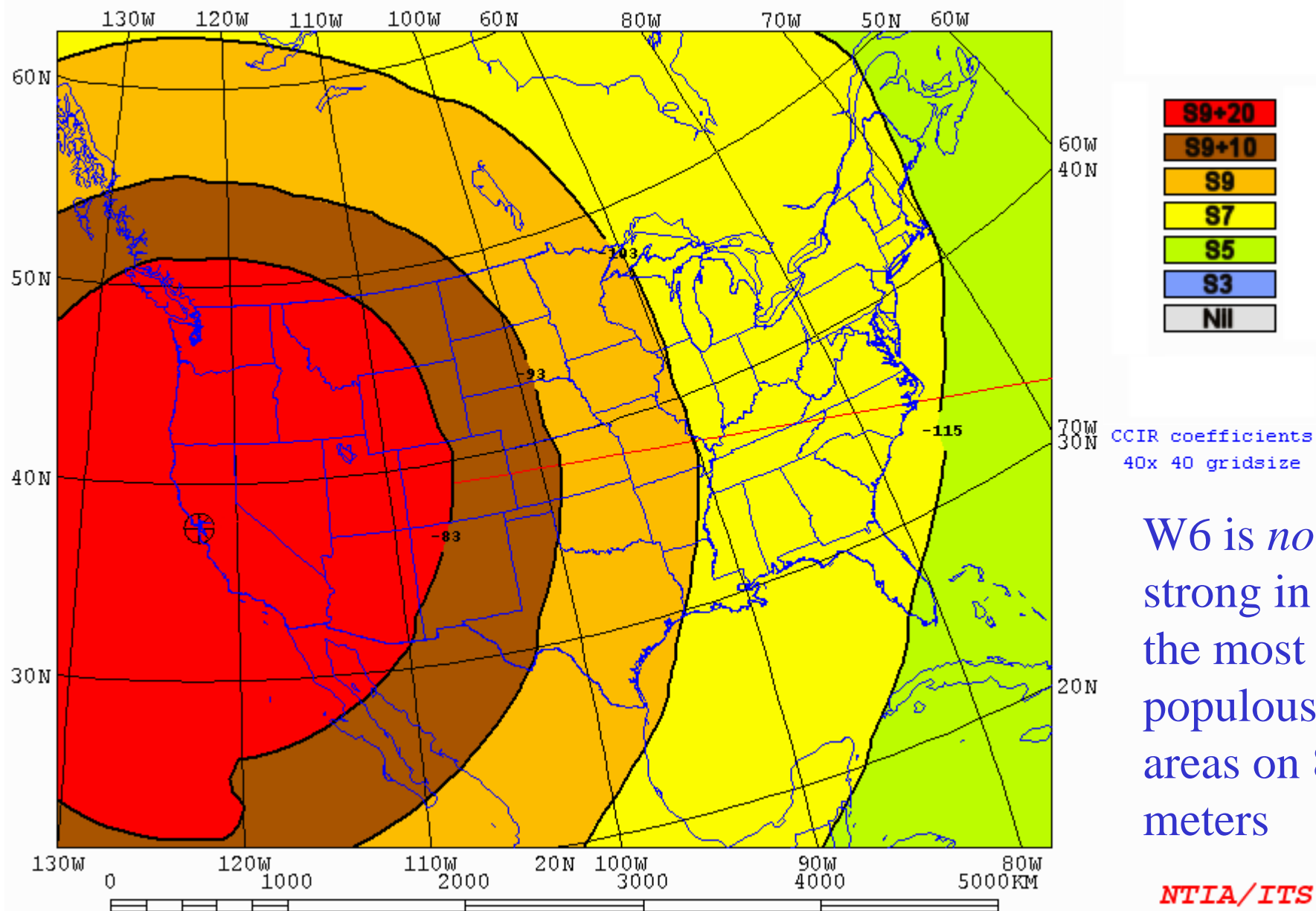
These maps were created using *VOAAREA*, part of the *VOACAP* (Voice of America Coverage Analysis Program) software suite.

SAN FRANCISCO [Dipole @ 7] 1.5kW 80deg 02ut 3.800MHz Nov 10ssn

SDBW

Tx location to grid of Rx

AREADATA\DEFAULT\SF4.V19

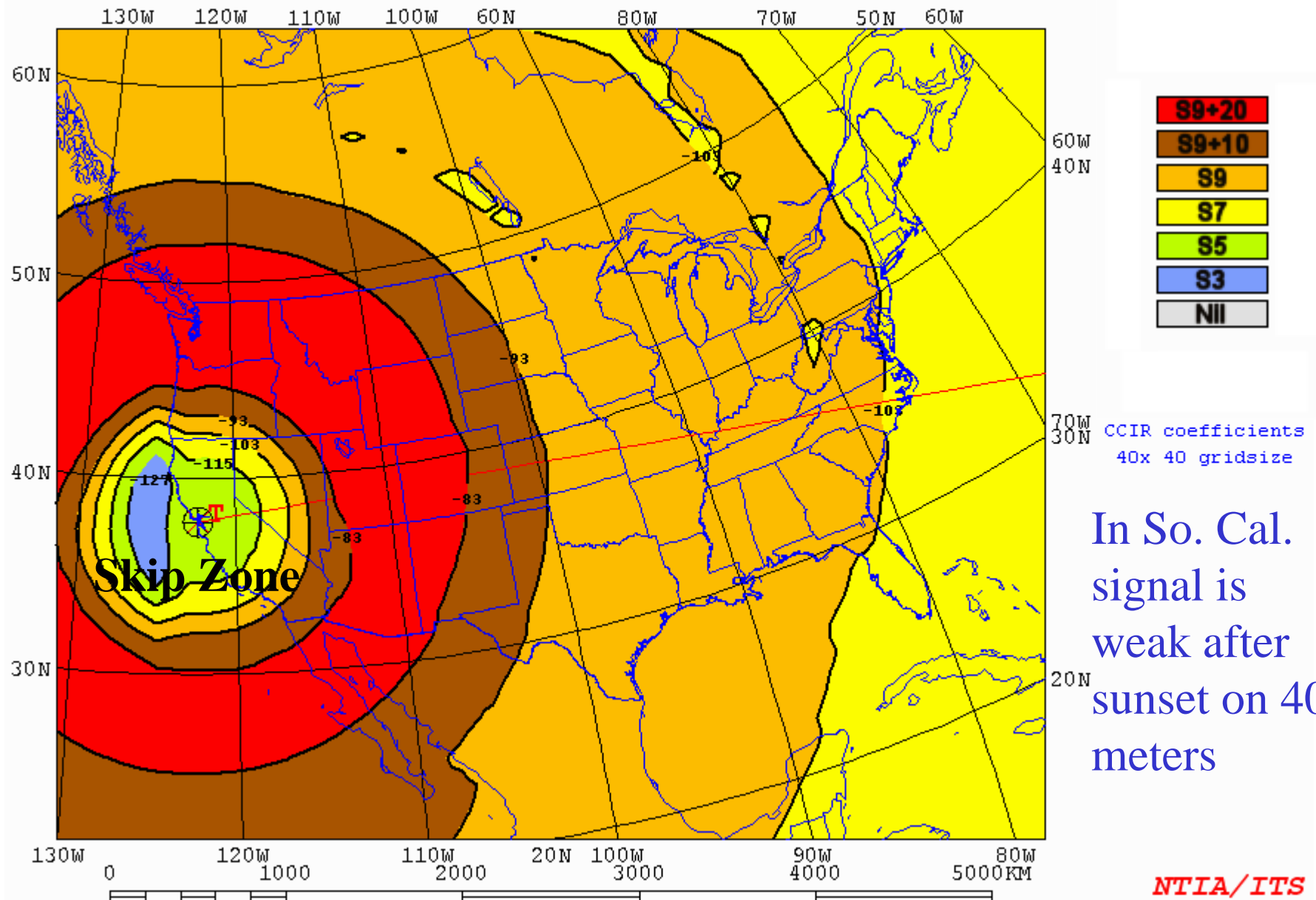


SAN FRANCISCO [Dipole @ 7] 1.5kW 80deg 02ut 7.200MHz Nov 10ssn

SDBW

Tx location to grid of Rx

AREADATA\DEFAULT\SF7.V19



In So. Cal.
signal is
weak after
sunset on 40
meters

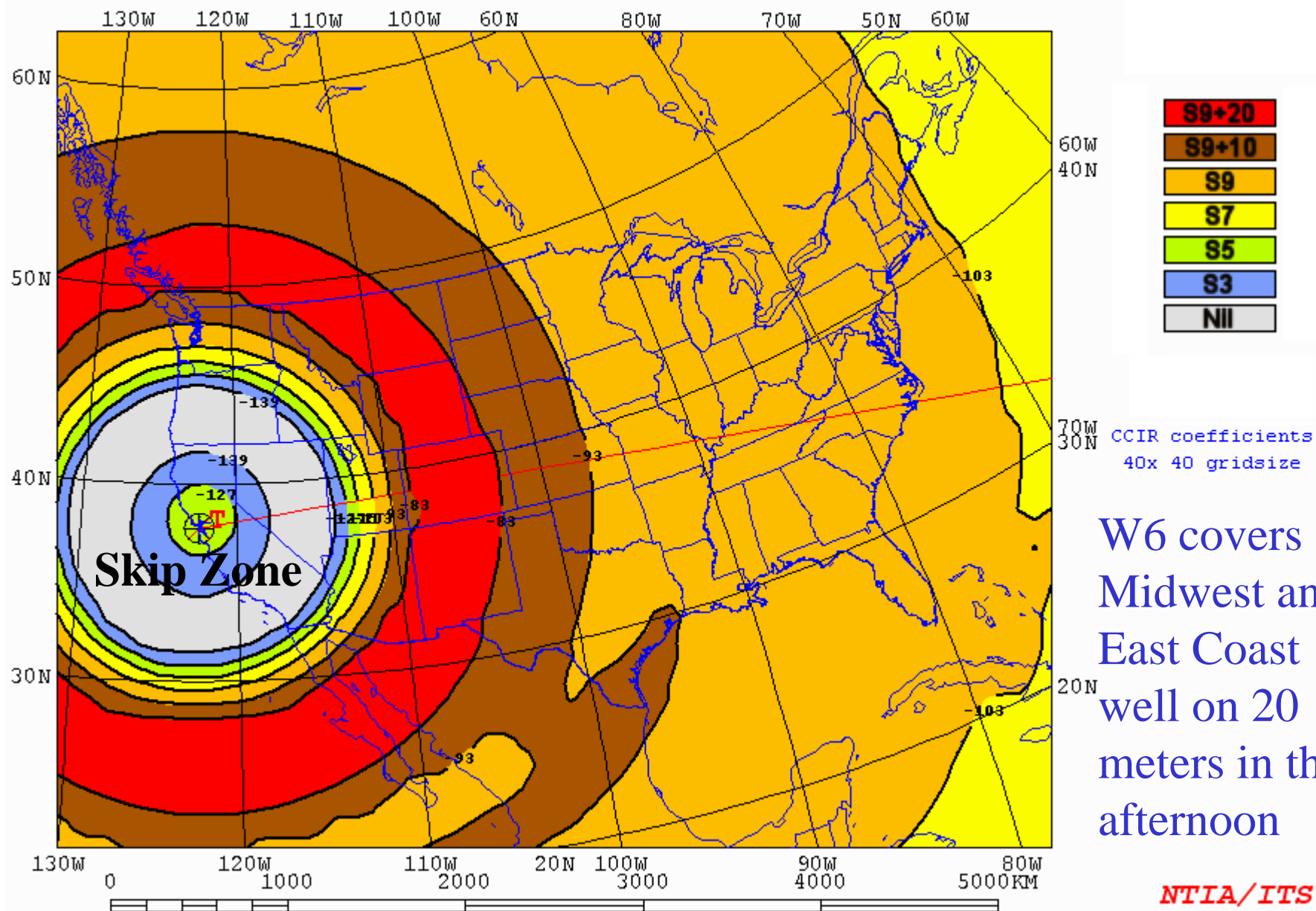
NTIA/ITS

SAN FRANCISCO [3-el Yagi] 1.5kW 80deg 22ut 14.200MHz Nov 10ssn

SDBW

Tx location to grid of Rx

AREADATA\DEFAULT\SF14.V15



W6 covers
Midwest and
East Coast
well on 20
meters in the
afternoon

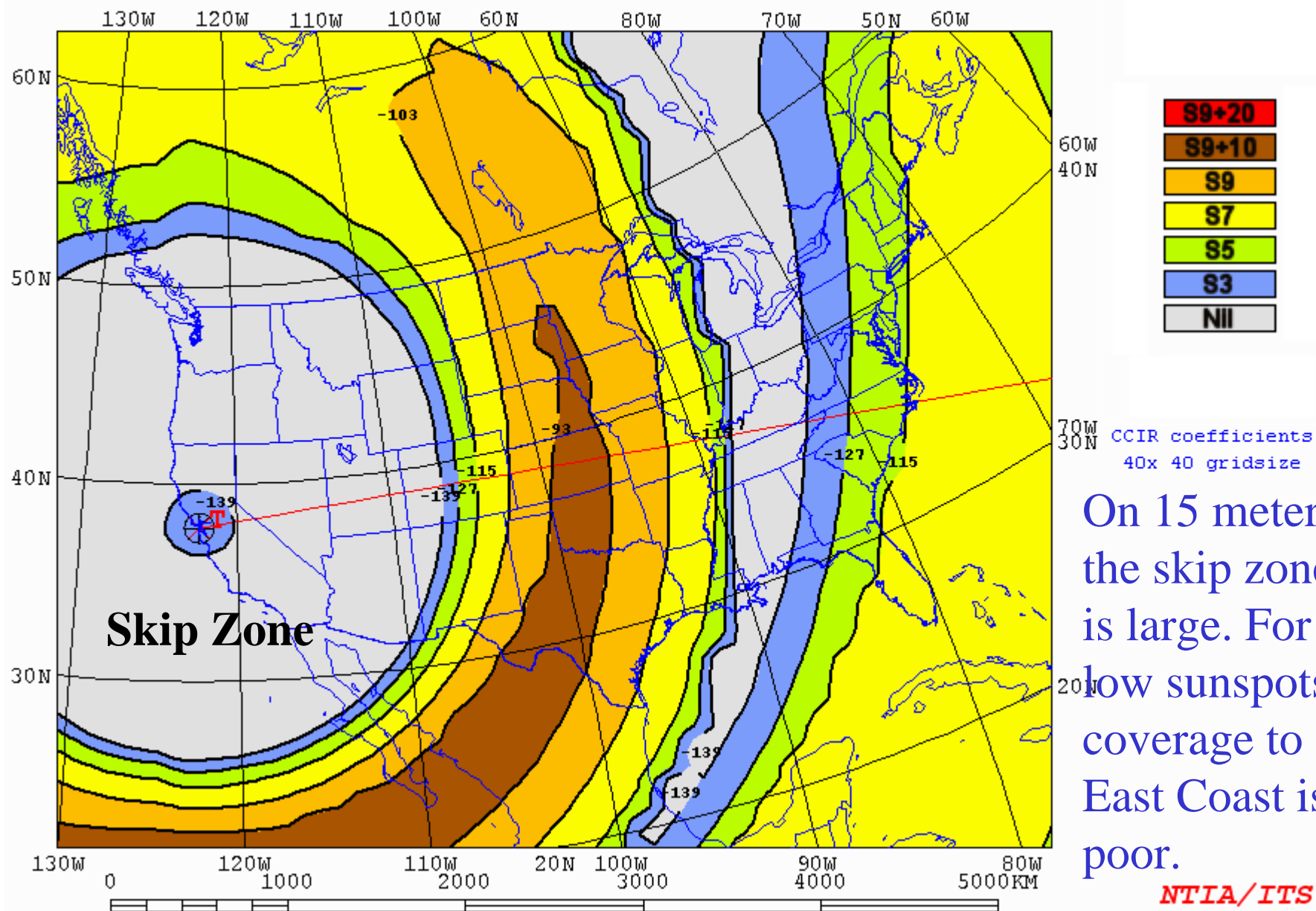
NTIA/ITS

SAN FRANCISCO [3-el Yagi] 1.5kW 80deg 18ut 21.200MHz Nov 10ssn

SDBW

Tx location to grid of Rx

AREADATA\DEFAULT\SF21.V11



“Exotic” Propagation Modes

- The HF propagation predictions just described are for either “short path” or “long path.”

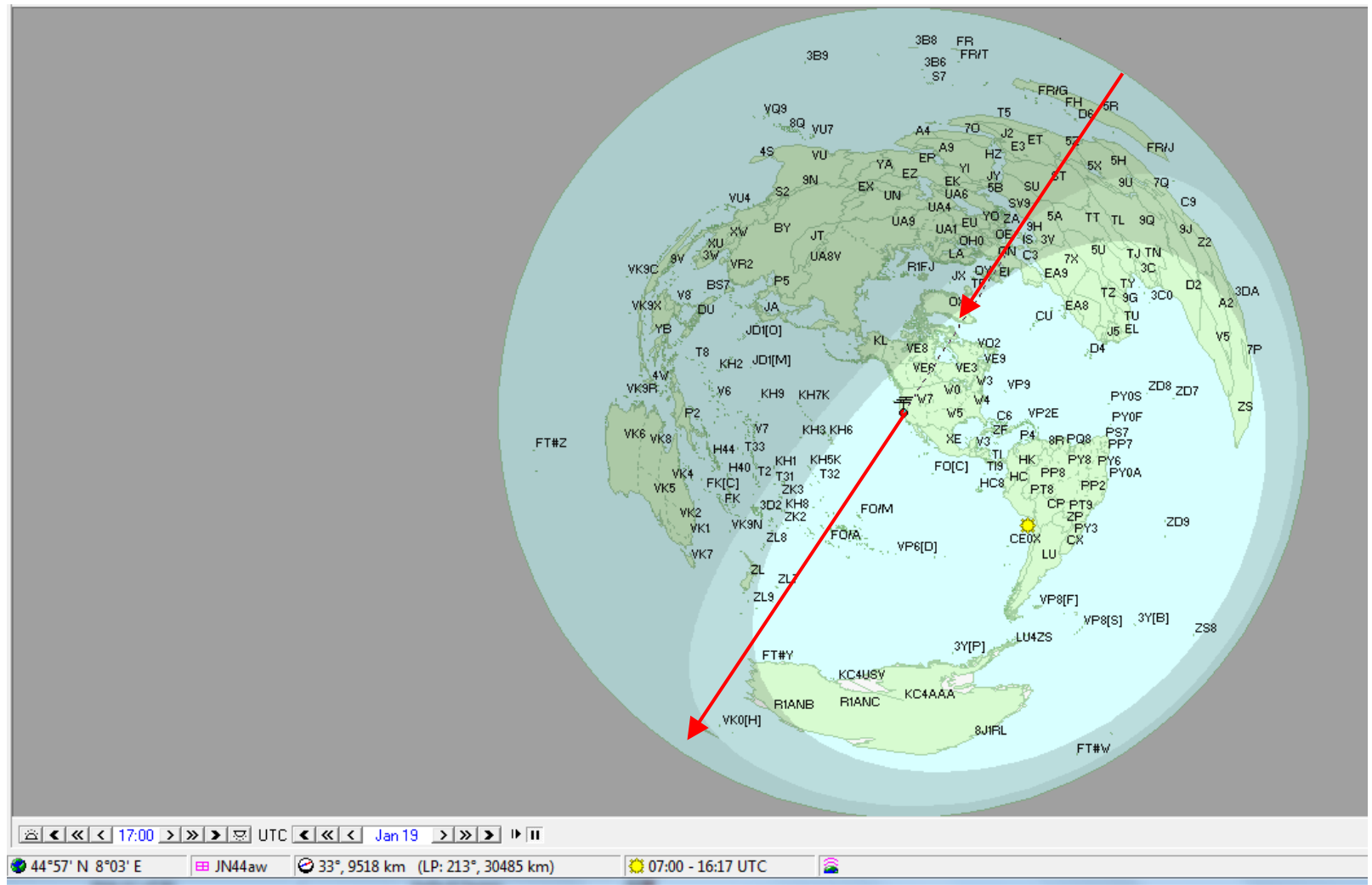
N6BV Propagation Predictions

10 Meters: Jan., CA (San Francisco), for SSN = High, Sigs in S-Units. (c) 2010 Dean Straw, N6BV

Zone	UTC -->																							
	00	01	02	03	04	05	06	07	08	09	10	11	12	13	14	15	16	17	18	19	20	21	22	23
KL7 = 01	9+	7	2*	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	8	9+	9+	9+	9+	9+
VO2 = 02	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	2*	5	9	8	6	1	-
W6 = 03	5	6	6	7	7	7	7	7	7	7	7	7	7	7	7	7	7	5	5	5	5	5	5	5
W9 = 04	4	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	7	9	9+	9+	9+	9+	9
W3 = 05	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	3*	3	3	2	-	-
XE1 = 06	9	4	3	-	-	-	-	-	-	-	-	-	-	-	-	3	9+	9+	9+	9+	9+	9+	9+	9+
TI = 07	5*	4*	-	-	-	-	-	-	-	-	-	-	-	-	-	-	7	9	9	9	9	7	7	5
VP2 = 08	5*	6*	2*	-	-	-	-	-	-	-	-	-	-	-	-	-	9	9	9	9	9	9	9	5*
P4 = 09	5*	5*	1*	-	-	-	-	-	-	-	-	-	-	-	-	1	9	9	9	9	9	9	9	5*
HC = 10	6*	5*	-	-	-	-	-	-	-	-	-	-	-	-	-	3	9	9	9	9	9	9	9	9
PY1 = 11	9	2	1	-	-	-	-	-	-	-	-	-	-	-	-	6	8	7	8	7	9	9	9	9
CE = 12	9	5	3	1	-	-	-	-	-	-	-	-	-	-	-	5	8	8	8	8	8	9	9	9
LU = 13	9	5	3	-	-	-	-	-	-	-	-	-	-	-	-	6	8	8	8	8	8	9	9	9
G = 14	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	6*	4*	2*	1*	-	-	-
I = 15	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-
UA3 = 16	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-
UN = 17	-	-	-	1*	-	-	-	-	-	-	-	-	-	-	-	-	2*	-	-	-	-	-	-	-
UA9 = 18	-	-	-	1*	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-
UA0 = 19	7	1	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	4
4X = 20	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	1*	7*	4*	2*	2*	2*	-	-
HZ = 21	-	-	-	-	2*	-	-	-	-	-	-	-	-	-	-	-	5*	3*	1*	-	-	-	-	-
VU = 22	-	-	-	1*	-	-	-	-	-	-	-	-	-	-	-	-	5*	3*	2*	-	-	-	-	-
JT = 23	1	6	2*	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-
VR2 = 24	8	7	1	1	-	-	-	-	-	-	-	-	-	-	-	4*	-	-	-	-	-	-	-	-
JA1 = 25	6	3	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	5
HS = 26	1	6	-	-	-	-	-	-	-	-	-	-	-	-	-	4*	6*	2*	1*	-	-	-	-	-
DU = 27	8	8	3	1	-	-	-	-	-	-	-	-	-	-	-	3*	1*	-	-	-	-	-	-	8
YB = 28	6	7	1	-	-	-	-	-	-	-	-	-	-	-	-	2*	1*	1*	-	-	-	-	-	-
VK6 = 29	-	1	5	3	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-
VK3 = 30	6	7	7	6	2	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	1	1	5
KH6 = 31	1	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	2	4	2	1	2
KH8 = 32	4	6	8	8	5	-	-	-	-	-	-	-	-	-	-	-	-	-	5	9	7	1	1*	1
CN = 33	2*	1*	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	7	4*	3*	2*	3*	4*	3*
SU = 34	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	1*	7*	3*	2*	1*	1*	-	-
6W = 35	3*	1*	1*	-	-	-	-	-	-	-	-	-	-	-	-	-	7	9	9	8	4	3*	3*	4*
D2 = 36	6	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	7	9	9	5	6	8	8	9
5Z = 37	3	-	-	-	2*	-	-	-	-	-	-	-	-	-	-	-	-	5*	8	9	8	8	7	6
ZS6 = 38	4	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	7	8	8	7	7	7	6	5
FR = 39	-	-	-	-	3*	-	-	-	-	-	-	-	-	-	-	-	-	-	3	8	8	6	4	-
FJL = 40	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-
Zone	00	01	02	03	04	05	06	07	08	09	10	11	12	13	14	15	16	17	18	19	20	21	22	23
* = Longpath																								

UTC --> * = Longpath
Expected signal levels using 1500 W and 14 dBi isotropic antennas.

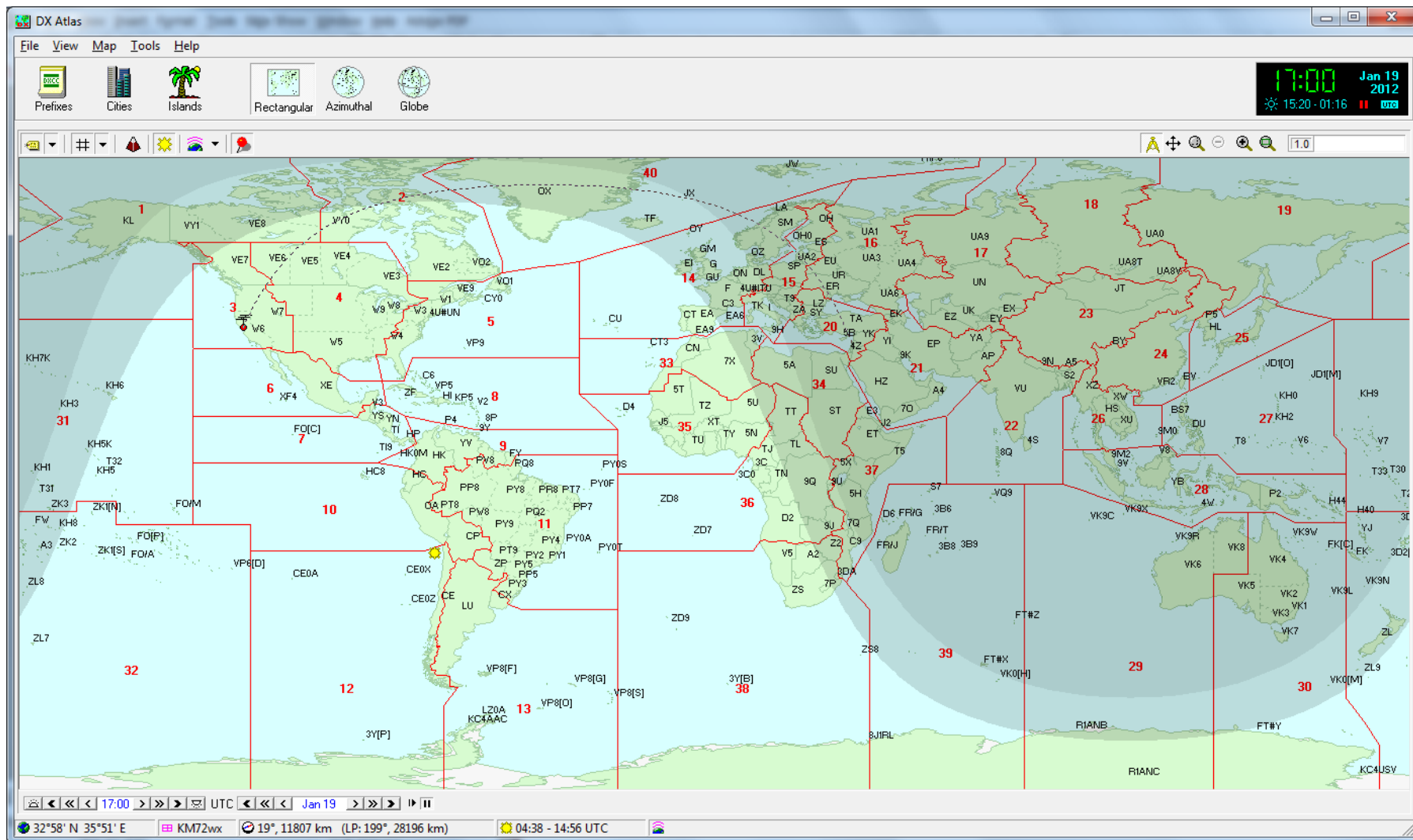
Long path to So. Europe/No. Africa on 10 meters.



33° short-
path to Italy

213° long-
path to Italy

DX Atlas with
azimuthal projection



19° short-path to Israel

199° long-path to Israel

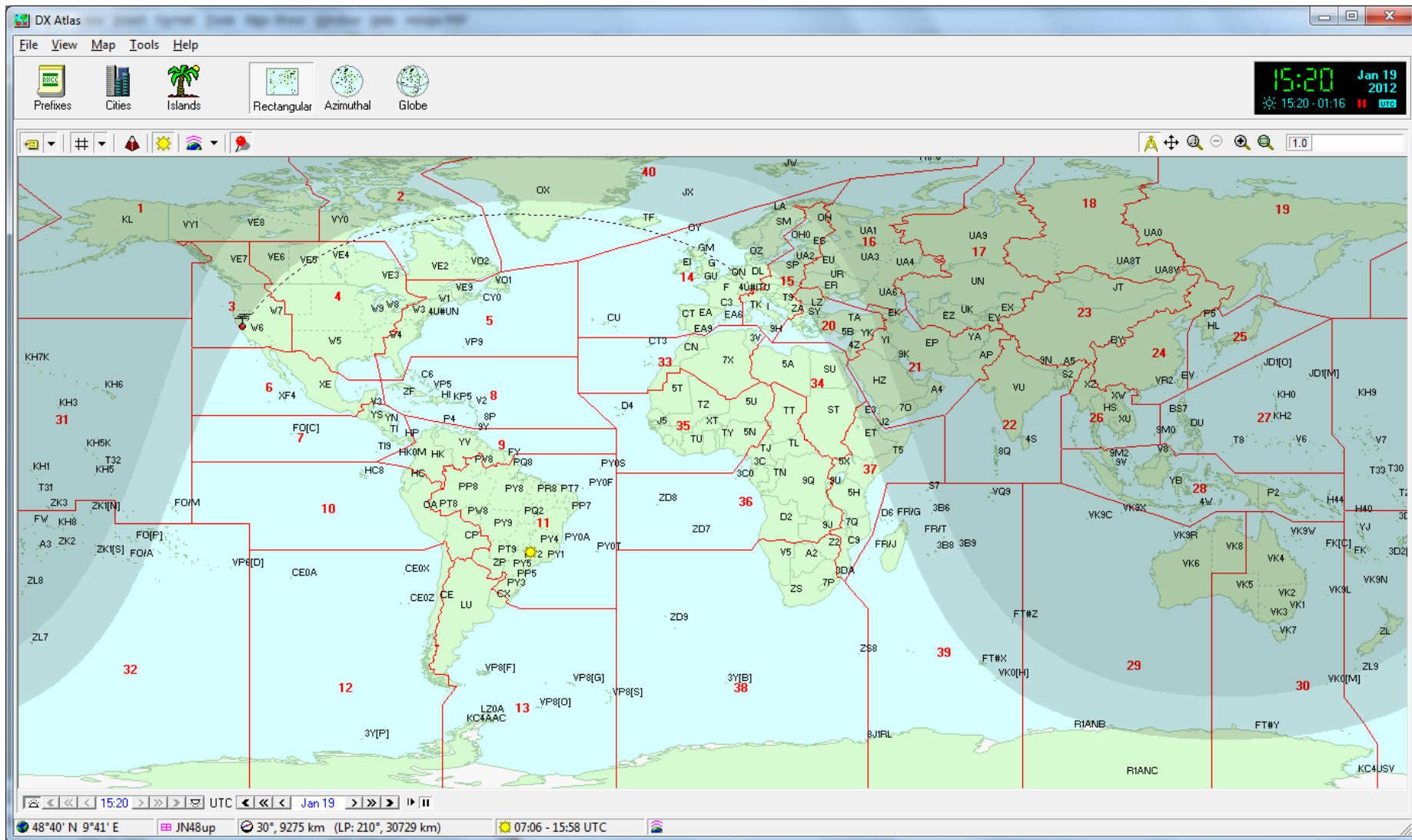
DX Atlas, with
Mercator projection

“Exotic” Propagation Modes

- The HF propagation predictions just described are for either “short path” or “long path.”

Long-path openings on the lower bands can be very short — even only a minute long on 160! Thus, predictions good for a whole hour may easily miss such transient openings.

Grayline information, however, can be very useful on the lower bands.



30° short-path
to Austria

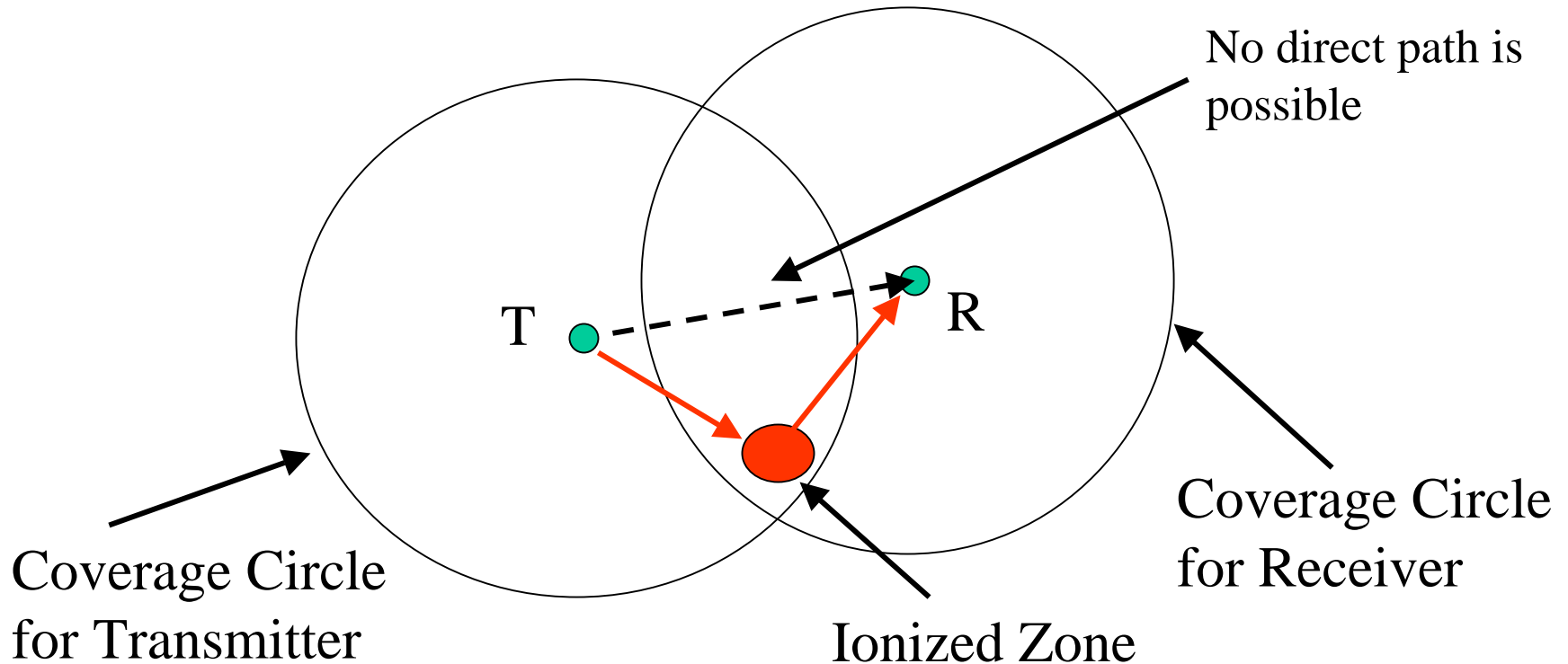
210° long-path to
Austria on 80 meters

DX Atlas, with
Mercator projection

“Exotic” Propagation Modes

- The HF propagation predictions just described are for either “short path” or “long path.”
- Sidescatter (skew) paths – especially when sunspots are low.

Sidescatter (Skew) Propagation



A common 15-meter skew-path opening is from W6 to Europe, with both stations pointing their beams at CT3. This occurs shortly before the morning direct short-path opening.

“Exotic” Propagation Modes

- The HF propagation predictions just described are for either “short path” or “long path.”
- Sidescatter (skew) paths – especially when sunspots are low.
- Backscatter paths – close-in stations.

Backscatter Paths

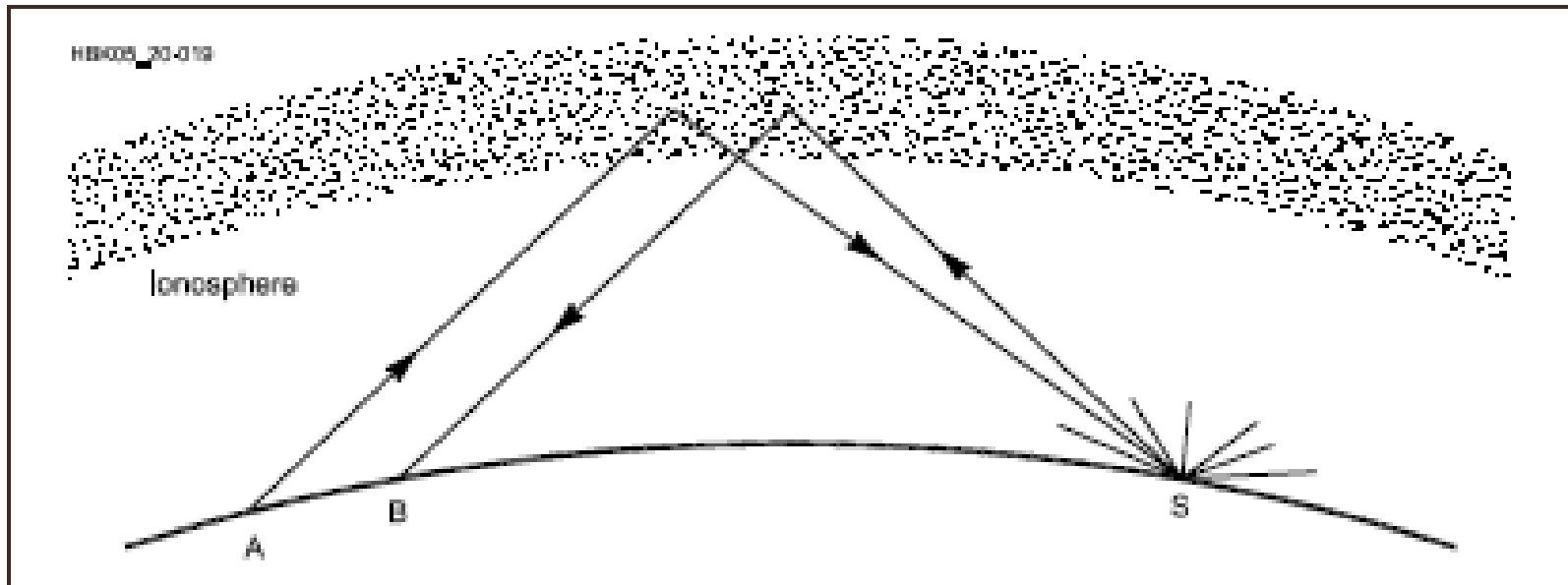


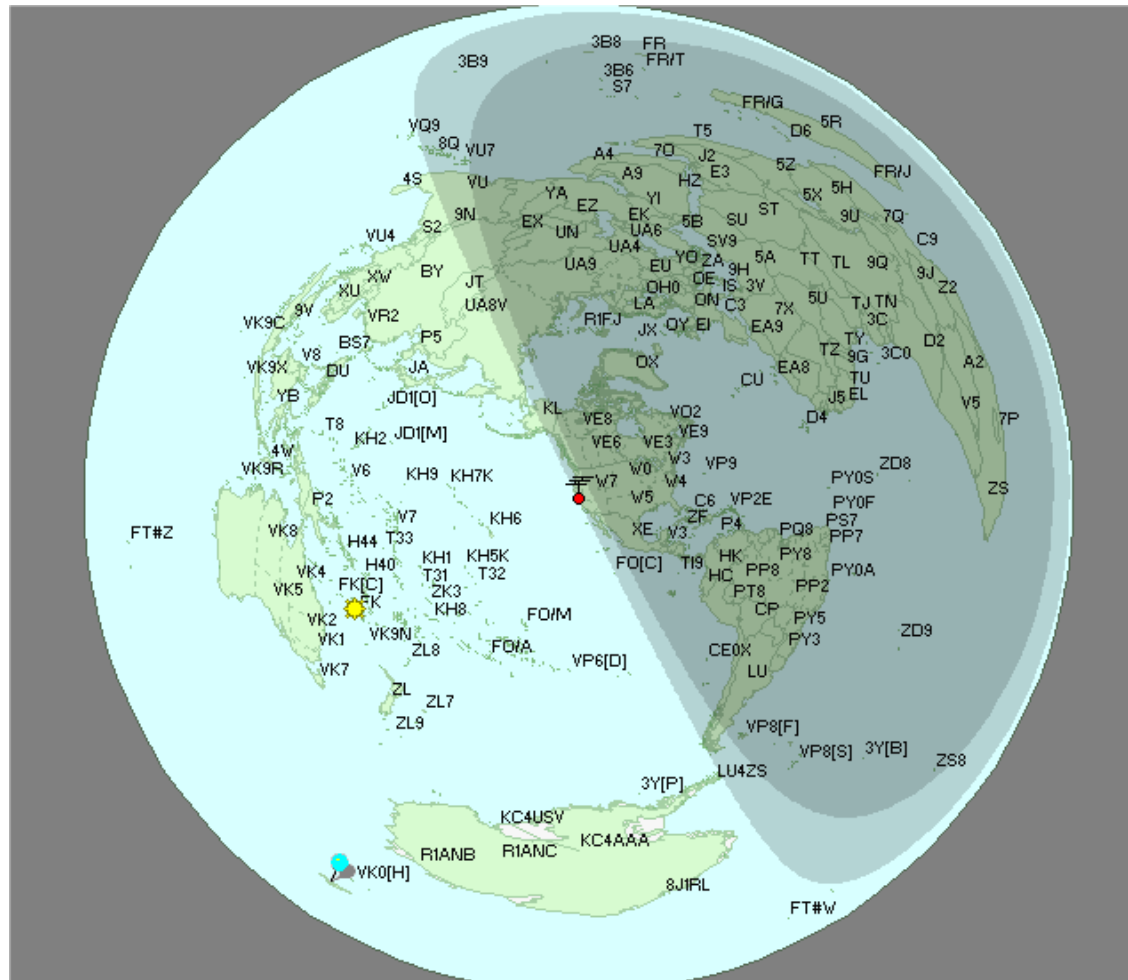
Fig 20.19—Schematic of a simple backscatter path. Stations A and B are too close to make contact via normal F-layer ionospheric refraction. Signals scattered back from a distant point on the Earth's surface (S), often the ocean, may be accessible to both and create a backscatter circuit.

This allows us to work locals on 15 or 10 meters in the SS.

“Exotic” Propagation Modes

- The HF propagation predictions just described are for either “short path” or “long path.”
- Sidescatter (skew) paths – especially when sunspots are low.
- Backscatter paths – close-in stations.
- Grayline paths – along the sunset/sunrise “terminator” on the lower bands.

“Exotic” Propagation Modes



Grayline short-path propagation: W6 (sunset) to VU2 (sunrise), late November on 40 meters.

“Exotic” Propagation Modes

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- Grayline paths – along the sunset/sunrise “terminator” on the lower bands.
- Transequatorial propagation.

“Exotic” Propagation Modes

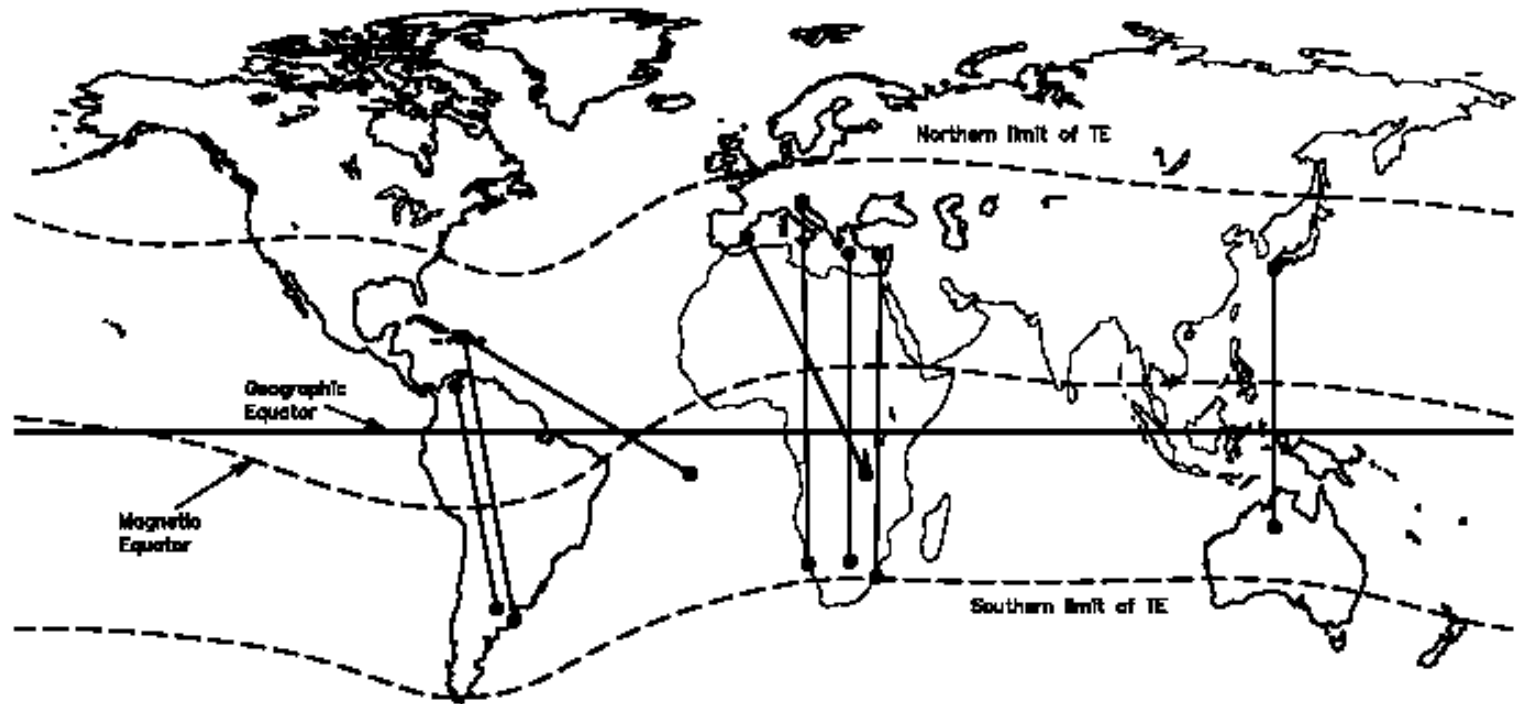


Fig 8—Transequatorial spread-F propagation takes place between stations equidistant across the geomagnetic equator. Distances up to 8000 km (5000 miles) are possible on 28 through 432 MHz. Note that the geomagnetic equator is considerably south of the geographic equator in the Western Hemisphere. (Figure courtesy of *The ARRL Handbook*.)

Transequatorial propagation on 10 meters.

“Exotic” Propagation Modes

- The HF propagation predictions just described are for either “short path” or “long path.”
- Sidescatter (skew) paths – especially when sunspots are low.
- Backscatter paths – close-in stations.
- Grayline paths – along the sunset/sunrise “terminator” on the lower bands.
- Transequatorial propagation.
- Sporadic-E.

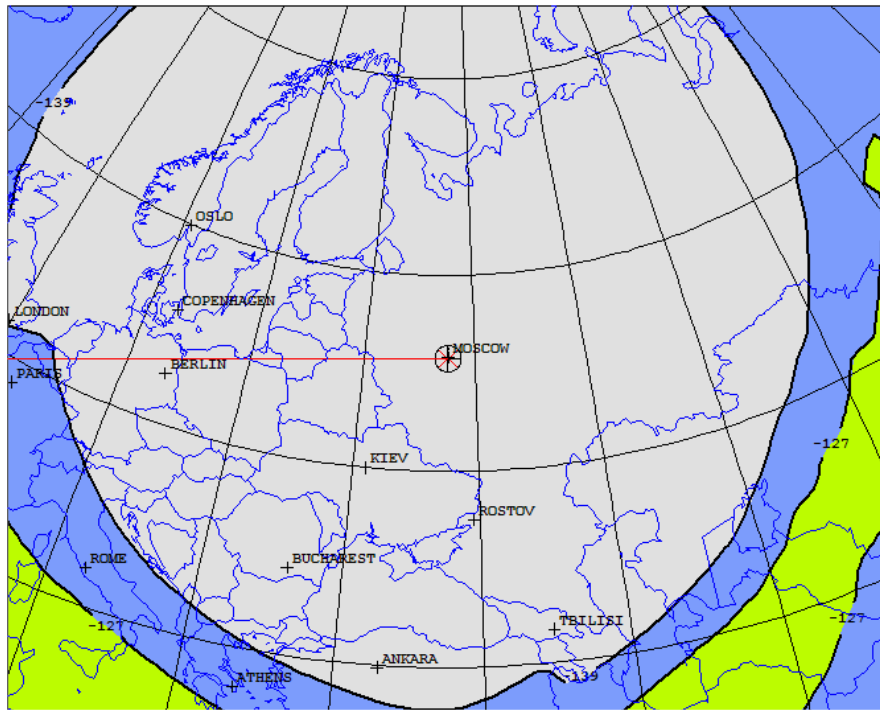
“Exotic” Propagation Modes

- HF Sporadic-E propagation (E_s) – most common on 6 and 10 meters.

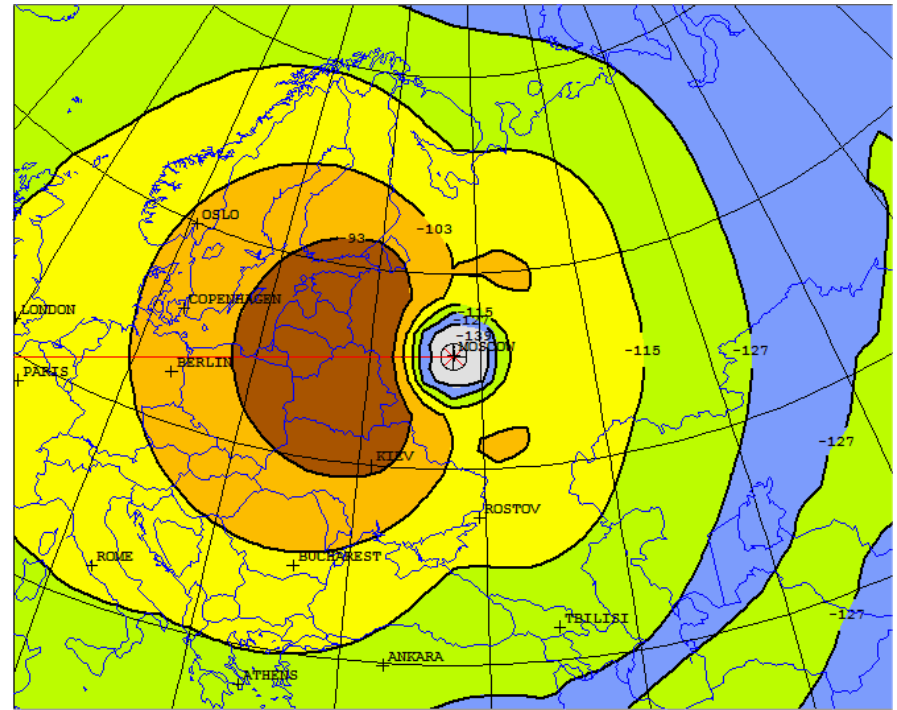
“Exotic” Propagation Modes

- HF Sporadic-E propagation (E_S) – most common on 6 and 10 meters.
- E_S goes even as low as 20 meters.

Side-by-Side Comparisons for Europe: July, 2010, 03 UTC (Sunrise, Zone 27), 20 Meters



Without Sporadic-E



With Sporadic-E

WRTC July 2010, Moscow

“Exotic” Propagation Modes

- HF Sporadic-E propagation (E_S) – most common on 6 and 10 meters.
- E_S goes even as low as 20 meters.
- E_S most common in December and June.

“Exotic” Propagation Modes

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- Sporadic-E “clouds” can be anywhere geographically.

“Exotic” Propagation Modes

- HF Sporadic-E propagation (E_S) – most common on 6 and 10 meters.
- E_S goes even as low as 20 meters.
- E_S most common in December and June.
- Sporadic-E “clouds” can be anywhere geographically.
- E_S signals can be exceptionally loud!

Summary

- Ionospheric propagation of HF signals is *very* complicated.

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- But that just makes it more interesting and challenging for an operator! (Where do I turn the beam — short-path, long-path, skew-path, grayline?)

Summary

- Ionospheric propagation of HF signals is *very* complicated.
- But that just makes it more interesting and challenging for an operator! (Where do I turn the beam — short-path, long-path, skew-path, grayline?)
- There are propagation-prediction programs that can help, but there's nothing like actually getting on the air and hearing what's coming in — before the big guns get there!



And Where Can You Get the N6BV Propagation Predictions?

- The exclusive distributor is *Radio-Ware* (also known as *Radio Bookstore*).

<http://www.radio-ware.com/>

- The price is \$30.
- Also see my webinar “Tactical Use of Propagation Predictions for HF Contesting” at:

<http://nccc.cc/webinars.html>

